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# Long-time convergence of an adaptive biasing force method: Variance reduction by Helmholtz projection 

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#### Abstract

In this paper, we propose an improvement of the adaptive biasing force (ABF) method, by projecting the estimated mean force onto a gradient. We show on some numerical examples that the variance of the approximated mean force is reduced using this technique, which makes the algorithm more efficient than the standard ABF method. The associated stochastic process satisfies a nonlinear stochastic differential equation. Using entropy techniques, we prove exponential convergence to the stationary state of this stochastic process.


Keywords. Adaptive biasing force; Helmholtz projection; Free energy; Variance reduction.

## 1. Introduction

### 1.1. The model

Let us consider the Boltzmann-Gibbs measure :

$$
\begin{equation*}
\mu(d x)=Z_{\mu}^{-1} \mathrm{e}^{-\beta V(x)} d x, \tag{1.1}
\end{equation*}
$$

where $x \in \mathcal{D}^{N}$ denotes the position of $N$ particles in $\mathcal{D}$. The space $\mathcal{D}$ is called the configuration space. One should think of $\mathcal{D}$ as a subset of $\mathbb{R}^{n}$, or the $n$-dimensional torus $\mathbb{T}^{n}$ (where $\mathbb{T}=\mathbb{R} / \mathbb{Z}$ denotes the one dimensional torus). The potential energy function $V: \mathcal{D} \longrightarrow \mathbb{R}$ associates with the positions of the particles $x \in \mathcal{D}$ its energy $V(x)$. In addition, $Z_{\mu}=\int_{\mathcal{D}} \mathrm{e}^{-\beta V(x)} d x$ (assumed to be finite) is the normalization constant and $\beta=1 /\left(k_{B} T\right)$ is proportional to the inverse of the temperature $T, k_{B}$ being the Boltzmann constant.

The probability measure $\mu$ is the equilibrium measure sampled by the particles in the canonical statistical ensemble. A typical dynamics that can be used to sample this measure is the Overdamped Langevin Dynamics:

$$
\begin{equation*}
d X_{t}=-\nabla V\left(X_{t}\right) d t+\sqrt{\frac{2}{\beta}} d W_{t} \tag{1.2}
\end{equation*}
$$

where $X_{t} \in \mathcal{D}^{N}$ and $W_{t}$ is a $N n$-dimensional standard Brownian motion. Under loose assumptions on $V$, the dynamics $\left(X_{t}\right)_{t \geq 0}$ is ergodic with respect to the measure $\mu$, which means: for any smooth test function $\varphi$,

$$
\begin{equation*}
\lim _{T \rightarrow+\infty} \frac{1}{T} \int_{0}^{T} \varphi\left(X_{t}\right) d t=\int \varphi d \mu \tag{1.3}
\end{equation*}
$$

i.e. trajectory averages converge to canonical averages.

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### 1.2. Metastability, reaction coordinate and free energy

In many cases of interest, there exists regions of the configuration space where the dynamics (1.2) remains trapped for a long time, and jumps only occasionally to another region, where it again remains trapped for a long time. This typically occurs when there exist high probability regions separated by very low probability areas. The regions where the process $\left(X_{t}\right)_{t \geq 0}$ remains trapped for very long times are called metastable.

Because of the metastability, trajectorial averages (1.3) converge very slowly to their ergodic limit. Many methods have been proposed to overcome this difficulty, and we concentrate here on the Adaptive Biasing Force (denoted ABF) method (see [5, 7]). In order to introduce the ABF method, we need another ingredient: a reaction coordinate (also known as an order parameter), $\xi=\left(\xi_{1}, \ldots, \xi_{m}\right): \mathcal{D} \longrightarrow$ $\mathbb{R}^{m}, \xi(x)=z$, where $m<n N$. Typically, in (1.2), the time-scale for the dynamics on $\xi\left(X_{t}\right)$ is longer than the time-scale for the dynamics on $X_{t}$ due to the metastable states, so that $\xi$ can be understood as a function such that $\xi\left(X_{t}\right)$ is in some sense a slow variable compared to $X_{t}$. In some sense, $\xi$ describes the metastable states of the dynamics associated with the potential $V$. For a given configuration $x$, $\xi(x)$ represents some macroscopic information. For example, it could represent angles or bond lengths in a protein, positions of defects in a material, etc. In any case, it is meant to be a function with values in a small dimensional space (i.e. $m \leq 4$ ), since otherwise, it is difficult to approximate accurately the associated free energy which is a scalar function defined on the range of $\xi$ (see equation (1.5) below). The choice of a "good" reaction coordinate is a highly debatable subject in the literature. One aim of the mathematical analysis conducted here or in previous papers (see for example [9]) is to quantify the efficiency of free energy biasing methods once a reaction coordinate has been chosen.

The image of the measure $\mu$ by $\xi$ is defined by

$$
\begin{equation*}
\xi * \mu:=\exp (-\beta A(z)) d z \tag{1.4}
\end{equation*}
$$

where $A$ is the so-called free energy associated with the reaction coordinate $\xi$. By the co-area formula (see [9], Appendix A), the following formula for the free energy can then be obtained: up to an additive constant,

$$
\begin{equation*}
A(z)=-\beta^{-1} \ln \left(Z_{\Sigma_{z}}\right) \tag{1.5}
\end{equation*}
$$

where $Z_{\Sigma_{z}}=\int_{\Sigma_{z}} \mathrm{e}^{-\beta V(x)} \delta_{\xi(x)-z}(d x)$, the submanifold $\Sigma_{z}$ is defined by

$$
\Sigma_{z}=\left\{x=\left(x_{1}, \ldots, x_{n}\right) \in \mathcal{D} \mid \xi(x)=z\right\}
$$

and $\delta_{\xi(x)-z}(d x)$ represents a measure with support $\Sigma_{z}$, such that $\delta_{\xi(x)-z}(d x) d z=d x$ (for further details on delta measures, we refer to [10], Section 3.2.1). We assume henceforth that $\xi$ and $V$ are such that $Z_{\Sigma_{z}}<\infty$, for all $z \in \mathbb{R}^{m}$.

The idea of free energy biasing methods, such as the adaptive biasing force method (see [5, 7]), or the Wang-Landau algorithm (see [14]), is that, if $\xi$ is well chosen, the dynamics associated with $V-A \circ \xi$ is less metastable than the dynamics associated with $V$. Indeed, from the definition of the free energy (1.4), for any compact subspace $\mathcal{M} \subset \mathbb{R}^{m}$, the image of $\tilde{Z}^{-1} \mathrm{e}^{(-\beta(V-A \circ \xi)(x))} 1_{\xi(x) \in \mathcal{M}}$ by $\xi$ is the uniform law $\frac{1_{\mathcal{M}}}{|\mathcal{M}|}$, where $\tilde{Z}=\int_{Z_{\Sigma_{z}}} \mathrm{e}^{(-\beta(V-A \circ \xi)(x))} 1_{\xi(x) \in \mathcal{M}} d x$ and $|\mathcal{M}|$ denotes the Lebesgue measure of $\mathcal{M}$. The uniform law is typically easier to sample than the original measure $\xi * \mu$. Therefore, if the function $\xi$ is well chosen (i.e. if the dynamics in the direction orthogonal to $\xi$ is not too metastable), the free energy can be used as a biasing potential to accelerate the sampling of the dynamics (see [9]). The difficulty is of course that the free energy $A$ is unknown and difficult to approximate using the original dynamics (1.2) because of metastability. Actually, in many practical cases, it is the quantity of interest that one would like to approximate by molecular dynamics simulations (see [4, 10]). The principle of adaptive biasing methods is thus to directly approximate $A$ (or its gradient) on the fly in order to bias the dynamics and to reduce the metastable features of the original dynamics (1.2). Methods

## Projected ABF method

which directly approximate $A$ are called Adaptive Biasing Potential methods whereas methods which compute an estimate of $\nabla A$ are called Adaptive Biasing Force methods. Here, we focus on the latter. For a comparison of these two approaches, we refer for example to [8] and Section 5.1.1.5 in [10].

### 1.3. Adaptive biasing force method (ABF)

In order to introduce the ABF method, we need a formula for the derivatives of $A$. The so called mean force $\nabla A(z)$, can be obtained from (1.5) as (see [10], Section 3.2.2):

$$
\begin{equation*}
\nabla A(z)=\int_{\Sigma_{z}} f(x) d \mu_{\Sigma_{z}}, \tag{1.6}
\end{equation*}
$$

where $d \mu_{\Sigma_{z}}$ is the probability measure $\mu$ conditioned to a fixed value $z$ of the reaction coordinate:

$$
\begin{equation*}
d \mu_{\Sigma_{z}}=Z_{\Sigma_{z}}^{-1} \mathrm{e}^{-\beta V(x)} \delta_{\xi(x)-z}(d x), \tag{1.7}
\end{equation*}
$$

and $f=\left(f_{1}, \ldots, f_{m}\right)$ is the so-called local mean force defined by

$$
\begin{equation*}
f_{i}=\sum_{j=1}^{m}\left(G^{-1}\right)_{i, j} \nabla \xi_{j} . \nabla V-\beta^{-1} \operatorname{div}\left(\sum_{j=1}^{m}\left(G^{-1}\right)_{i, j} \nabla \xi_{j}\right), \tag{1.8}
\end{equation*}
$$

where $G=\left(G_{i, j}\right)_{i, j=1, \ldots, m}$, has components $G_{i, j}=\nabla \xi_{i} \cdot \nabla \xi_{j}$. This can be rewritten in terms of conditional expectation as: for a random variable $X$ with law $\mu$ (defined by (1.1)),

$$
\begin{equation*}
\nabla A(z)=\mathbb{E}(f(X) \mid \xi(x)=z) . \tag{1.9}
\end{equation*}
$$

In the literature, what is called the mean force is sometimes the negative of $\nabla A$. We stick here to the terminology used in the references $[8,9,10]$. We are now in a position to introduce the standard adaptive biasing force ( ABF ) technique, applied to the overdamped Langevin dynamics (1.2) for a given initial condition $X_{0} \in \mathbb{R}^{N n}$, independent of $\left(W_{t}\right)_{t \geq 0}$ :

$$
\left\{\begin{align*}
d X_{t} & =-\left(\nabla V-\sum_{i=1}^{m} F_{t}^{i} \circ \xi \nabla \xi_{i}+\nabla\left(W_{c} \circ \xi\right)\right)\left(X_{t}\right) d t+\sqrt{2 \beta^{-1}} d W_{t},  \tag{1.10}\\
F_{t}^{i}(z) & =\mathbb{E}\left[f_{i}\left(X_{t}\right) \mid \xi\left(X_{t}\right)=z\right], i=1, \ldots, m,
\end{align*}\right.
$$

where $f$ is defined in (1.8). One should emphasize that the expectation in the Equation (1.10), unlike that of (1.9), is over all realizations of the stochastic differential equation. Compared with the original dynamics (1.2), two modifications have been made to obtain the ABF dynamics (1.10):
(1) First and more importantly, the force $\sum_{i=1}^{m} F_{t}^{i} \circ \xi \nabla \xi_{i}$ has been added to the original force $-\nabla V$. At time $t, F_{t}$ approximates $\nabla A$ defined in (1.6).
(2) Second, a potential $W_{c} \circ \xi$ has been added. This is actually only needed in the case when $\xi$ lives in an unbounded domain. In this case, a so-called confining potential $W_{c}$ is introduced so that the law of $\xi\left(X_{t}\right)$ admits a longtime limit $Z_{W_{c}}^{-1} \mathrm{e}^{-\beta W_{c}(z)} d z$ (see Remark 3.3 at the end of Section(3.2), where $Z_{W_{c}}=\int_{\operatorname{Ran}(\xi)} \mathrm{e}^{-\beta W_{c}}$ is assumed to be finite. When $\xi$ is living in a compact subspace of $\mathbb{R}^{m}$, there is no need to introduce such a potential and the law of $\xi\left(X_{t}\right)$ converges exponentially fast to the uniform law on the compact subspace (as explained in Section 1.2 and Section(3.2). Typically, $W_{c}$ is zero in a chosen compact subspace $\mathcal{M}$ of $\mathbb{R}^{m}$

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and is harmonic outside $\mathcal{M}$. For example, in dimension two, suppose that $\xi=\left(\xi_{1}, \xi_{2}\right)$ and $\mathcal{M}=\left[\xi_{\text {min }}, \xi_{\text {max }}\right] \times\left[\xi_{\text {min }}, \xi_{\text {max }}\right]$, then $W_{c}$ can be defined as:

$$
\begin{equation*}
W_{c}\left(z_{1}, z_{2}\right)=\sum_{i=1}^{2} 1_{z_{i} \geq \xi_{\text {max }}}\left(z_{i}-\xi_{\text {max }}\right)^{2}+\sum_{i=1}^{2} 1_{z_{i} \leq \xi_{\text {min }}}\left(z_{i}-\xi_{\text {min }}\right)^{2} . \tag{1.11}
\end{equation*}
$$

It is proven in [9] that, under appropriate assumptions, $F_{t}$ converges exponentially fast to $\nabla A$. In addition, for well chosen $\xi$, the convergence to equilibrium for (1.10) is much quicker than for (1.2). This can be quantified using entropy estimates and Logarithmic Sobolev Inequalities, see [9].

Notice that even though $F_{t}$ converges to a gradient $(\nabla A)$, there is no reason why $F_{t}$ would be a gradient at time $t$. In this paper, we propose an alternative method, where we approximate $\nabla A$, at any time $t$, by a gradient denoted $\nabla A_{t}$. The gradient $\nabla A_{t}$ is defined as the Helmholtz projection of $F_{t}$. One could expect improvements compared to the original ABF method since the variance of $\nabla A_{t}$ is then smaller than the variance of $F_{t}$ (since $A_{t}$ is a scalar function). Reducing the variance is important since the conditional expectation in (1.10) is approximated by empirical averages in practice.

### 1.4. Projected adaptive biasing force method (PABF)

A natural algorithm to reconstruct $A_{t}$ from $F_{t}$, consists in solving the following Poisson problem:

$$
\begin{equation*}
\Delta A_{t}=\operatorname{div} F_{t} \quad \text { on } \mathcal{M}, \tag{1.12}
\end{equation*}
$$

with appropriate boundary conditions depending on the choice of $\xi$ and $\mathcal{M}$. More precisely, if $\xi$ is periodic and $\mathcal{M}$ is the torus $\mathbb{T}^{m}$, then we are working with periodic boundary conditions. If $\xi$ is with values in $\mathbb{R}^{m}$ and $\mathcal{M}$ is a bounded subset of $\mathbb{R}^{m}$, then Neumann boundary conditions are needed (see Remark 3.17 at the end of Section(3.3.2). To discretize this Poisson problem, standard methods such as finite difference methods, finite element methods, spectral methods (with, for example, a fast Fourier transform to solve the associated linear system efficiently) can be used. Note that (1.12) is the Euler-Lagrange equation associated with the minimization problem:

$$
\begin{equation*}
A_{t}=\underset{g \in H^{1}(\mathcal{M}) / \mathbb{R}}{\operatorname{argmin}} \int_{\mathcal{M}}\left|\nabla g-F_{t}\right|^{2}, \tag{1.13}
\end{equation*}
$$

where $H^{1}(\mathcal{M}) / \mathbb{R}=\left\{g \in H^{1}(\mathcal{M}) \mid \int_{\mathcal{M}} g=0\right\}$ denotes the subspace of $H^{1}(\mathcal{M})$ of zero average functions. In the following we denote by

$$
\begin{equation*}
\mathcal{P}\left(F_{t}\right)=\nabla A_{t} \tag{1.14}
\end{equation*}
$$

the projection of $F_{t}$ onto a gradient. In view of (1.13), $A_{t}$ can be interpreted as the function such that its gradient is the closest to $F_{t}$. Solving (1.12) amounts to computing the so-called Helmholtz-Hodge decomposition of the vector field $F_{t}$ as (see [6], Section 3):

$$
\begin{equation*}
F_{t}=\nabla A_{t}+R_{t}, \quad \text { on } \mathcal{M}, \tag{1.15}
\end{equation*}
$$

where $R_{t}$ is a divergence free vector field.
Finally, the projected $A B F$ dynamics we propose to study is the following nonlinear stochastic differential equation:

$$
\left\{\begin{align*}
d X_{t} & =-\nabla\left(V-A_{t} \circ \xi+W_{c} \circ \xi\right)\left(X_{t}\right) d t+\sqrt{2 \beta^{-1}} d W_{t},  \tag{1.16}\\
\Delta A_{t} & =\operatorname{div} F_{t} \quad \text { on } \mathcal{M}, \text { with appropriate boundary conditions, } \\
F_{t}^{i}(z) & =\mathbb{E}\left[f_{i}\left(X_{t}\right) \mid \xi\left(X_{t}\right)=z\right], i=1, \ldots, m .
\end{align*}\right.
$$

Compared with the standard ABF dynamics (1.10), the only modification is that the mean force $F_{t}$ is replaced by $\nabla A_{t}$, which is meant to be an approximation of $\nabla A$ at time $t$.

## Projected ABF method

The main theoretical result of this paper is a proof that $A_{t}$ converges exponentially fast to the free energy $A$ : we actually prove this result in a simple setting, namely when the domain is the torus, $\xi\left(x_{1}, \ldots, x_{n}\right)=\left(x_{1}, x_{2}\right)$ and for a slightly modified version of (1.16), see Section 3 for more details, and Remarks 3.17 and 3.18 for discussions about possible extensions. Moreover, we illustrate numerically this result on a prototypical example. From a numerical point of view, the interest of the method is that the variance of the norm of the projected estimated mean force (i.e. $\nabla A_{t}$ ) is smaller than the variance of the norm of the estimated mean force (i.e. $F_{t}$ ). We observe numerically that this variance reduction enables a faster convergence to equilibrium for PABF compared with the original ABF.

The paper is organized as follows. Section 2 is devoted to a numerical illustration of the practical value of the projected ABF compared to the standard ABF approach. In Section 3, the longtime convergence of the projected ABF method is proven. Finally, the proofs of the results presented in Section 3 are provided in Section 4.

## 2. Numerical experiments

### 2.1. Presentation of the model

We consider a system composed of $N$ particles $\left(q_{i}\right)_{0 \leq i \leq N-1}$ in a two-dimensional periodic box of side length $L$. Among these particles, three particles (numbered 0,1 and 2 in the following) are designated to form a trimer, while the others are solvent particles. In this model, a trimer is a molecule composed of three identical particles linked together by two bonds (see Figure(2.1).

### 2.1.1. Potential functions

All particles, except the three particles forming the trimer, interact through the purely repulsive WCA pair potential, which is the Lennard-Jones (LJ) potential truncated at the LJ potential minimum:

$$
V_{W C A}(d)=\left\{\begin{array}{cl}
\varepsilon+4 \varepsilon\left[\left(\frac{\sigma}{d}\right)^{12}-\left(\frac{\sigma}{d}\right)^{6}\right] & \text { if } d \leq d_{0} \\
0 & \text { if } d \geq d_{0}
\end{array}\right.
$$

where $d$ denotes the distance between two particles, $\varepsilon$ and $\sigma$ are two positive parameters and $d_{0}=$ $2^{1 / 6} \sigma$.

A particle of the solvent and a particle of the trimer also interact through the potential $V_{W C A}$. The interaction potential between two consecutive particles of the trimer ( $q_{0} / q_{1}$ or $q_{1} / q_{2}$ ) is the double-well potential (see Figure(2.2):

$$
\begin{equation*}
V_{S}(d)=h\left[1-\frac{\left(d-d_{1}-\omega\right)^{2}}{\omega^{2}}\right]^{2} \tag{2.1}
\end{equation*}
$$

where $d_{1}, h$ and $\omega$ are positive parameters.
The potential $V_{S}$ has two energy minima. The first one, at $d=d_{1}$, corresponds to the compact bond. The second one, at $d=d_{1}+2 \omega$, corresponds to the stretched bond. The height of the energy barrier separating the two states is $h$.

In addition, the interaction potential between $q_{0}$ and $q_{2}$ is a Lennard-Jones potential:

$$
V_{L J}(d)=4 \varepsilon^{\prime}\left[\left(\frac{\sigma^{\prime}}{d}\right)^{12}-\left(\frac{\sigma^{\prime}}{d}\right)^{6}\right]
$$

where $\varepsilon^{\prime}$ and $\sigma^{\prime}$ are two positive parameters.
Finally, the three particles of the trimer also interact through the following potential function on the angle $\theta$ formed by the vectors $\overrightarrow{q_{1} q_{0}}$ and $\overrightarrow{q_{1} q_{2}}$ :

$$
V_{\theta_{0}}(\theta)=\frac{k_{\theta}}{2}\left(\cos (\theta)-\cos \left(\theta_{0}\right)\right)^{2},
$$

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where $\theta_{0}$ is the equilibrium angle and $k_{\theta}$ is the angular stiffness. Figure 2.1 presents a schematic view of the system.


Figure 2.1. Trimer $\left(q_{0}, q_{1}, q_{2}\right)$ Left: compact state; Center: mixed state; Right: stretched state.

The total energy of the system is therefore, for $x=\left(q_{0}, \ldots, q_{N-1}\right) \in(L \mathbb{T})^{2 N}$ :

$$
\begin{aligned}
V(x)= & \sum_{3 \leq i<j \leq N-1} V_{W C A}\left(\left|q_{i}-q_{j}\right|\right)+\sum_{i=0}^{2} \sum_{j=3}^{N-1} V_{W C A}\left(\left|q_{i}-q_{j}\right|\right) \\
& +\sum_{i=0}^{1} V_{S}\left(\left|q_{i}-q_{i+1}\right|\right)+V_{L J}\left(\left|q_{0}-q_{2}\right|\right)+V_{\theta_{0}}(\theta),
\end{aligned}
$$



Figure 2.2. Double-well potential (2.1), with $d_{1}=2^{1 / 6}, \omega=2$ and $h=2$.

### 2.1.2. Reaction coordinate and physical parameters

The reaction coordinate describes the transition from compact to stretched state in each bond. It is the normalised bond length of each bond of the trimer molecule. More precisely, the reaction coordinate is $\xi=\left(\xi_{1}, \xi_{2}\right)$, with $\xi_{1}(x)=\frac{\left|q_{0}-q_{1}\right|-d_{1}}{2 \omega}$ and $\xi_{2}(x)=\frac{\left|q_{1}-q_{2}\right|-d_{1}}{2 \omega}$, where $x=\left(q_{0}, \ldots, q_{N-1}\right) \in(L \mathbb{T})^{2 N}$. For $i=1,2$, the value $\xi_{i}=0$ refers to the compact state (i.e. $d=d_{1}$ ) and the value $\xi_{i}=1$ corresponds to the stretched state (i.e. $d=d_{1}+2 \omega$ ).

We apply ABF and PABF dynamics to the trimer problem described above. The inverse temperature is $\beta=1$, we use $N=100$ particles $(N-3$ solvent particles and the trimer) and the box side length is $L=15$. The parameters describing the WCA and the Lennard-Jones interactions are set to $\sigma=1$, $\varepsilon=1, \sigma^{\prime}=1, \varepsilon^{\prime}=0.1, d_{0}=2^{1 / 6}, d_{1}=2^{1 / 6}$ and the additional parameters for the trimer are $\omega=2$ and $h=2$. The parameters describing the angle potential are: $\theta_{0}$ such that $\cos \left(\theta_{0}\right)=1 / 3$ and $k_{\theta}=1$ (we refer to [12], Section 10.4.2, for the choice of such parameters). The initial condition on the trimer is as follows: Both bonds $q_{0} q_{1}$ and $q_{1} q_{2}$ are in compact state, which means that the distance between $q_{0}$ and $q_{1}$ and the distance between $q_{1}$ and $q_{2}$ are equal to $d_{0}$. Moreover, the initial bond angle is $\theta_{0}$.

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### 2.1.3. Numerical methods and numerical parameters

Standard and projected ABF methods are used with $N_{\text {replicas }}=100$ replicas of the system evolving according to the overdamped Langevin dynamics discretized with a time-step $\Delta t=2.5 \times 10^{-4}$. The reaction coordinate space of interest is taken of the form $\mathcal{M}=\left[\xi_{\min }, \xi_{\max }\right] \times\left[\xi_{\min }, \xi_{\max }\right]$, where $\xi_{\min }=-0.2$ and $\xi_{\max }=1.2$. The space $\mathcal{M}$ is discretized into $N_{b i n s} \times N_{b i n s}=50 \times 50=2500$ bins of equal sizes and $\delta=\delta_{x}=\delta_{y}=\frac{\xi_{\text {max }}-\xi_{\text {min }}}{N_{\text {bins }}}=0.028$ denotes the size of each bin along both axes.

To implement the ABF and PABF methods, one needs to approximate

$$
F_{t}^{i}\left(z_{1}, z_{2}\right)=\mathbb{E}\left[f_{i}\left(X_{t}\right) \mid\left(\xi_{1}\left(X_{t}\right), \xi_{2}\left(X_{t}\right)\right)=\left(z_{1}, z_{2}\right)\right], i=1,2
$$

The mean force $F_{t}$ is estimated in each bin as a combination of plain trajectorial averages and averages over replicas. It is calculated at each time as an average of the local mean force in the bin over the total number of visits in this bin. More precisely, at time $t$ and for $l=1,2$, the value of the mean force in the $(i, j)^{t h}$ bin is:

$$
\begin{equation*}
F_{t}^{l}(i, j)=\frac{\sum_{t^{\prime} \leq t} \sum_{k=1}^{N_{\text {replicas }}} f_{l}\left(X_{t^{\prime}}^{k}\right) 1_{\left\{i n d x\left(\xi\left(X_{t^{\prime}}^{k}\right)\right)=(i, j)\right\}}}{\left.\sum_{t^{\prime} \leq t} \sum_{k=1}^{N_{\text {replicas }}} 1_{\{i n d x}\left(\xi\left(X_{t^{\prime}}^{k}\right)\right)=(i, j)\right\}} \tag{2.2}
\end{equation*}
$$

where $X_{t^{\prime}}^{k}$ denotes the position at time $t$ of the $k$-th replica, $f=\left(f_{1}, f_{2}\right)$ is defined in (1.8) and $\operatorname{indx}(\xi(x))$ denotes the number of the bin where $\xi(x)$ lives, i.e.

$$
\operatorname{indx}(z)=\left(\left[\frac{z_{1}-\xi_{\min }}{\delta}\right]_{+},\left[\frac{z_{2}-\xi_{\min }}{\delta}\right]_{+}\right), \forall z=\left(z_{1}, z_{2}\right) \in \mathcal{M}
$$

If the components of the index function (i.e. indx) are either equal to -1 or to $N_{b i n s}$, it means that we are outside $\mathcal{M}$ and then the confining potential is non zero, and defined as:

$$
W_{c}\left(z_{1}, z_{2}\right)=\sum_{i=1}^{2}\left[1_{\left\{z_{i} \geq \xi_{\max }\right\}}\left(z_{i}-\xi_{\max }\right)^{2}+1_{\left\{z_{i} \leq \xi_{\min }\right\}}\left(z_{i}-\xi_{\min }\right)^{2}\right]
$$

To construct the PABF method, the solution to the following Poisson problem with Neumann boundary conditions is approximated:

$$
\begin{cases}\Delta A=\operatorname{div} F & \text { in } \mathcal{M}=\left[\xi_{\min }, \xi_{\max }\right] \times\left[\xi_{\min }, \xi_{\max }\right]  \tag{2.3}\\ \frac{\partial A}{\partial n}=F \cdot n & \text { on } \partial \mathcal{M}\end{cases}
$$

where $n$ denotes the unit normal outward to $\mathcal{M}$. The associated variational formulation is the following:

$$
\left\{\begin{array}{l}
\text { Find } A \in H^{1}(\mathcal{M}) / \mathbb{R} \text { such that } \\
\int_{\mathcal{M}} \nabla A \cdot \nabla v=\int_{\mathcal{M}} F \cdot \nabla v, \forall v \in H^{1}(\mathcal{M}) / \mathbb{R}
\end{array}\right.
$$

Problem (2.3) is solved using a finite element method of type $Q^{1}$ on the quadrilateral mesh defined above, with nodes $\left(z_{1}^{i}, z_{2}^{j}\right)$, where $z_{1}^{i}=\xi_{\text {min }}+i \delta$ and $z_{2}^{j}=\xi_{\text {min }}+j \delta$, for $i, j=0, . ., N_{\text {bins }}$. The space $\mathcal{M}$ is thus discretized into $N_{T}=N_{\text {bins }}^{2}$ squares, with $N_{s}=\left(N_{\text {bins }}+1\right)^{2}$ nodes.

In our numerical experiments, we do not try to optimize the computation of the solution to the Poisson problem (2.3). Notice that this Poisson problem is in small dimension (namely the number of reaction coordinates) and over a simple domain (the cartesian product of the domains of each reaction coordinates). Moreover, only the right-hand side of this problem changes from one time step to the other. Therefore very efficient numerical techniques can be used, such as fast Fourier transforms, finite differences methods over cartesian meshes by precomputing the LU decomposition of the matrix

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discretizing the Laplacian, etc. In practice, this is therefore negligible compared to the computation of the force in the dynamics.

### 2.2. Comparison of the methods

In this section, we compare results obtained with three different simulations: without ABF, with ABF and with projected ABF (PABF). First, it is observed numerically that both ABF methods overcome metastable states. Second, it is illustrated how PABF method reduces the variance of the estimated mean force compared to ABF method. As a consequence of this variance reduction, we observe that the convergence of $\nabla A_{t}$ to $\nabla A$ with the PABF method is faster than the convergence of $F_{t}$ to $\nabla A$ with the ABF method.

### 2.2.1. Metastability

To illustrate numerically the fact that ABF methods improve the sampling for metastable processes, we observe the variation, as a function of time, of the two metastable distances (i.e. the distance between $q_{0}$ and $q_{1}$, and the distance between $q_{1}$ and $q_{2}$ ). On Figures 2.3 and 2.4, the distance between $q_{1}$ and $q_{2}$ is plotted as a function of time for three dynamics: without ABF, ABF and PABF.

Both ABF methods allow to switch faster between the compact and stretched bonds and thus to better explore the set of configurations. Without adding the biasing term, the system remains trapped in a neighborhood of the first potential minimum (i.e. $d_{0} \simeq 1.12$ ) region for 20 units of time at least (see Figure(2.3), while when the biasing term is added in the dynamics, many jumps between the two local minima are observed (see Figure(2.4).


Figure 2.3. Without ABF.


Figure 2.4. Left: ABF; Right: PABF.

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### 2.2.2. Variance reduction

Since we use Monte-Carlo methods to approximate $F_{t}$ and $\nabla A_{t}$ (see Equation (2.2)), the variance is an important quantity to assess the quality of the result. The following general proposition shows that projection reduces the variance. It is written for a reaction coordinate with values in $\mathbb{T}^{2}$ but the generalization to other settings is straightforward.

Proposition 2.1. Let $F$ be a random function from $\mathbb{T}^{2}$ into $\mathbb{R}^{2}$ which belongs to $H\left(\operatorname{div}, \mathbb{T}^{2}\right)$, and denote by $\mathcal{P}$ the projection onto gradient vector fields defined by (1.14). Then, the variance of $\mathcal{P}(F)$ is smaller than the variance of $F$ :

$$
\int_{\mathbb{T}^{2}} \operatorname{Var}(\mathcal{P}(F)) \leq \int_{\mathbb{T}^{2}} \operatorname{Var}(F)
$$

where, for any vector field $F, \operatorname{Var}(F)=\mathbb{E}\left(|F|^{2}\right)-\mathbb{E}(|F|)^{2}$ and $|F|$ being the Euclidian norm.
Proof. Let $F$ be a random vector field of $H\left(\right.$ div, $\left.\mathbb{T}^{2}\right)$. Let us introduce $\mathcal{P}(F) \in H^{1}\left(\mathbb{T}^{2}\right) \times H^{1}\left(\mathbb{T}^{2}\right)$ its projection. Notice that by the linearity of the projection $\mathcal{P}(\mathbb{E}(F))=\mathbb{E}(\mathcal{P}(F))$. By definition of $\mathcal{P}(F)$, one gets:

$$
\int_{\mathbb{T}^{2}}(F-\mathcal{P}(F)) \cdot \nabla h=0, \forall h \in H^{1}\left(\mathbb{T}^{2}\right)
$$

Therefore, using the Pythagorean theorem and the fact that $\mathcal{P}(F)$ is a gradient,

$$
\int_{\mathbb{T}^{2}}|F|^{2}=\int_{\mathbb{T}^{2}}|F-\mathcal{P}(F)|^{2}+\int_{\mathbb{T}^{2}}|\mathcal{P}(F)|^{2}
$$

and

$$
\int_{\mathbb{T}^{2}}|F-\mathbb{E}(F)|^{2}=\int_{\mathbb{T}^{2}}|F-\mathbb{E}(F)-\mathcal{P}(F-\mathbb{E}(F))|^{2}+\int_{\mathbb{T}^{2}}|\mathcal{P}(F-\mathbb{E}(F))|^{2}
$$

Using the linearity of $\mathcal{P}$, we thus obtain

$$
\int_{\mathbb{T}^{2}} \operatorname{Var}(F)=\int_{\mathbb{T}^{2}} \operatorname{Var}(F-\mathcal{P}(F))+\int_{\mathbb{T}^{2}} \operatorname{Var}(\mathcal{P}(F))
$$

which concludes the proof.

We illustrate the improvement of the projected method in terms of the variances of the biasing forces by comparing $\int_{\mathcal{M}} \operatorname{Var}\left(\nabla A_{t}\right)=\int_{\mathcal{M}} \operatorname{Var}\left(\partial_{1} A_{t}\right)+\int_{\mathcal{M}} \operatorname{Var}\left(\partial_{2} A_{t}\right)$ (for the PABF method) with $\int_{\mathcal{M}} \operatorname{Var}\left(F_{t}\right)=\int_{\mathcal{M}} \operatorname{Var}\left(F_{t}^{1}\right)+\int_{\mathcal{M}} \operatorname{Var}\left(F_{t}^{2}\right)$ (for the ABF method). Figure 2.5 shows that the variance for the projected ABF method is smaller than for the standard ABF method. We have $N_{\text {bins }} \times N_{\text {bins }}=2500$ degrees of freedom for each term (i.e. $\partial_{1} A_{t}, \partial_{2} A_{t}, F_{t}^{1}$ and $F_{t}^{2}$ ). The variances are computed using 20 independent realizations as follows:

$$
\int_{\mathcal{M}} \operatorname{Var}\left(F_{t}^{1}\right) \simeq \frac{1}{2500} \sum_{i, j=1}^{50}\left[\frac{1}{20} \sum_{k=1}^{20} F_{t}^{1, k}\left(z_{1}^{i}, z_{2}^{j}\right)^{2}-\left(\frac{1}{20} \sum_{k=1}^{20} F_{t}^{1, k}\left(z_{1}^{i}, z_{2}^{j}\right)\right)^{2}\right]
$$

Note that four averages are involved in this formula: an average with respect to the space variable, an average over the 20 Monte-Carlo realizations, an average over replicas and a trajectorial average (the last two averages are more explicit in (2.2)). Since the variance of the biasing force is smaller with PABF, one may expect better convergence in time results. This will be investigated in Section 2.2.3 and 2.2.4.

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Figure 2.5. Variances as a function of time.

### 2.2.3. Free energy error

We now present, the variation, as a function of time, of the normalized averages $L^{2}$-distance between the real free energy (computed using a very long ABF simulation over $10^{6}$ time steps and 100 replicas) and the estimated one, in both cases: ABF and PABF methods. The average $L^{1}$-distance is approximated using 50 independent realizations. As can be seen in Figure 2.6, in both methods, the error decreases as time increases. Moreover, this error is always smaller for the projected ABF method than for the ABF method. At time 10, a good approximation of the free energy is obtained with PABF, while a similar result is only reached at time 25 with ABF.


Figure 2.6. Free energy error as a function of time.

### 2.2.4. Distribution

Another way to illustrate that the projected ABF method converges faster than the standard ABF method is to plot the density function $\psi^{\xi}$ as a function of time. The density $\psi^{\xi}$ is approximated using average over replicas and 50 independent realizations (see Figures 2.7(2.11).

It is observed that, for the projected ABF method, the state where both bonds are stretched is visited earlier (at time 5) than for the standard ABF method (at time 20). The convergence to uniform law along $\left(\xi_{1}, \xi_{2}\right)$ is faster with the projected ABF method.


Figure 2.7. At time 0.025. Left: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{2}$; Right: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{1}$.


Figure 2.8. At time 5. Left: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{2}$; Right: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{1}$.


Figure 2.9. At time 10. Left: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{2}$; Right: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{1}$.

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Figure 2.10. At time 20. Left: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{2} ;$ Right: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{1}$.


Figure 2.11. At time 25. Left: $\int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{2} ; \operatorname{Right}: \int \psi^{\xi}\left(z_{1}, z_{2}\right) d z_{1}$.

## 3. Longtime convergence of the projected ABF method

We study the longtime convergence in a simplified setting (see Remark 3.17 and 3.18 for possible extensions). We assume in this Section that $\mathcal{D}=\mathbb{T}^{n}$ and that $\xi(x)=\left(x_{1}, x_{2}\right)$. Then $\xi$ lives in the compact space $\mathcal{M}=\mathbb{T}^{2}$ and we therefore take $W_{c}=0$. The free energy can be written as:

$$
\begin{equation*}
A\left(x_{1}, x_{2}\right)=-\beta^{-1} \ln \left(Z_{\left(\Sigma_{\left(x_{1}, x_{2}\right)}\right)}\right) \tag{3.1}
\end{equation*}
$$

where $Z_{\Sigma_{\left(x_{1}, x_{2}\right)}}=\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \mathrm{e}^{-\beta V(x)} d x_{3} \ldots d x_{n}$ and $\Sigma_{\left(x_{1}, x_{2}\right)}=\left\{x_{1}, x_{2}\right\} \times \mathbb{T}^{n-2}$. The mean force becomes:

$$
\begin{equation*}
\nabla A\left(x_{1}, x_{2}\right)=\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} f(x) d \mu_{\left(x_{1}, x_{2}\right)}, \tag{3.2}
\end{equation*}
$$

where $f=\left(f_{1}, f_{2}\right)=\left(\partial_{1} V, \partial_{2} V\right)$ and the conditional probability measure $d \mu_{\Sigma_{\left(x_{1}, x_{2}\right)}}$ is:

$$
d \mu_{\Sigma_{\left(x_{1}, x_{2}\right)}}=Z_{\Sigma_{\left(x_{1}, x_{2}\right)}}^{-1} \mathrm{e}^{-\beta V} d x_{3} \ldots d x_{n} .
$$

Finally, the vector field $F_{t}\left(x_{1}, x_{2}\right)$ is $\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} f d \mu_{\left(x_{1}, x_{2}\right)}(t,$.$) , or equivalently:$

$$
F_{t}^{i}\left(x_{1}, x_{2}\right)=\mathbb{E}\left(\partial_{i} V\left(X_{t}\right) \mid \xi\left(X_{t}\right)=\left(x_{1}, x_{2}\right)\right), i=1,2 .
$$

### 3.1. Helmholtz projection

In section 3.1.1, a weighted Helmholtz-Hodge decomposition of $F_{t}$ is presented. In section 3.1.2, the associated minimization problem and projection operator are introduced.

Let us first fix some notations. For $x \in \mathbb{T}^{n}$ and $1 \leq i<j \leq n, x_{i}^{j}$ denotes the vector ( $x_{i}, x_{i+1}, \ldots, x_{j}$ ) and $d x_{i}^{j}$ denotes $d x_{i} d x_{i+1} \ldots d x_{j}$. Moreover, $\nabla_{x_{1}^{2}}, \operatorname{div}_{x_{1}^{2}}$ and $\Delta_{x_{1}^{2}}$ represent respectively the gradient,

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the divergence and the laplacian in dimension two for the first two variables $\left(x_{1}, x_{2}\right)$. Likewise, $\nabla_{x_{3}^{n}}=$ $\left(\partial_{3}, \ldots, \partial_{n}\right)^{T}$ represents the gradient vector starting from the third variable of $\mathbb{T}^{n}$.

### 3.1.1. Helmholtz decomposition

The space $\mathbb{T}^{2}$ is a bounded and connected space. For any smooth positive probability density function $\varphi: \mathbb{T}^{2} \rightarrow \mathbb{R}$, let us define the weighted Hilbert space: $L_{\varphi}^{2}\left(\mathbb{T}^{2}\right)=\left\{f: \mathbb{T}^{2} \rightarrow \mathbb{R}, \int_{\mathbb{T}^{2}}|f|^{2} \varphi<\infty\right\}$. Let us also introduce the Hilbert space $H_{\varphi}\left(\operatorname{div} ; \mathbb{T}^{2}\right)=\left\{g \in L_{\varphi}^{2}\left(\mathbb{T}^{2}\right) \times L_{\varphi}^{2}\left(\mathbb{T}^{2}\right), \operatorname{div}_{x_{1}^{2}}(g) \in L^{2}\left(\mathbb{T}^{2}\right)\right\}$. It is well-known that any vector field $F_{t}: \mathbb{T}^{2} \rightarrow \mathbb{R}^{2} \in H_{1}\left(\operatorname{div}, \mathbb{T}^{2}\right)$ can be written (see [6], Section 3 for example) as (Helmholtz decomposition): $F_{t}=\nabla_{x_{1}^{2}} A_{t}+R_{t}$, where $R_{t}$ is a divergence free vector field. We will need a generalization of the standard Helmholtz decomposition to the weighted Hilbert spaces $L_{\varphi}^{2}\left(\mathbb{T}^{2}\right)$ and $\left.H_{\varphi}\left(\operatorname{div} ; \mathbb{T}^{2}\right)\right)$ :

$$
\begin{equation*}
F_{t} \varphi=\nabla_{x_{1}^{2}}\left(A_{t}\right) \varphi+R_{t} \tag{3.3}
\end{equation*}
$$

s.t. $\operatorname{div}_{x_{1}^{2}}\left(R_{t}\right)=0$. This weighted Helmholtz decomposition is required to simplify calculations when studying the longtime convergence (see Remark 4.1 in Section 4.1 for more details). Recall the space: $H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}=\left\{g \in H^{1}\left(\mathbb{T}^{2}\right) \mid \int_{\mathbb{T}^{2}} g=0\right\}$. The function $A_{t}$ is then the solution to the following problem:

$$
\begin{equation*}
\int_{\mathbb{T}^{2}} \nabla_{x_{1}^{2}} A_{t} \cdot \nabla_{x_{1}^{2}} g \varphi=\int_{\mathbb{T}^{2}} F_{t} \cdot \nabla_{x_{1}^{2}} g \varphi, \quad \forall g \in H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R} \tag{3.4}
\end{equation*}
$$

which is the weak formulation of the Poisson problem:

$$
\begin{equation*}
\operatorname{div}_{x_{1}^{2}}\left(\nabla_{x_{1}^{2}} A_{t} \varphi(t, .)\right)=\operatorname{div}_{x_{1}^{2}}\left(F_{t} \varphi(t, .)\right) \tag{3.5}
\end{equation*}
$$

with periodic boundary conditions. Using standard arguments (Lax-Milgram theorem), it is straightforward to check that (3.4) admits a unique solution $A_{t} \in H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}$.

### 3.1.2. Minimization problem and projection onto a gradient

Proposition 3.1. Suppose that $F_{t} \in H_{\varphi}\left(\operatorname{div} ; \mathbb{T}^{2}\right)$. Then for any smooth positive probability density function $\varphi$, Equation (3.4) is the Euler Lagrange equation associated with the following minimization problem:

$$
\begin{equation*}
A_{t}=\min _{h \in H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}} \int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} h-F_{t}\right|^{2} \varphi=\min _{h \in H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}}\left\|\nabla_{x_{1}^{2}} h-F_{t}\right\|_{L_{\varphi}^{2}\left(\mathbb{T}^{2}\right)}^{2} \tag{3.6}
\end{equation*}
$$

Furthermore, $A_{t}$ belongs to $H^{2}\left(\mathbb{T}^{2}\right)$.
Proof. Let us introduce the application $I: H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R} \rightarrow \mathbb{R}_{+}$, defined by $I(g)=\left\|\nabla_{x_{1}^{2}} h-F_{t}\right\|_{L_{\varphi}^{2}\left(\mathbb{T}^{2}\right)}^{2}$. It is easy to prove that $I$ is $\alpha$-convex and coercive, i.e. $\lim _{\|g\|_{H^{1} \rightarrow+\infty}} I(g)=+\infty$. Thus $I$ admits a unique global minimum $A_{t} \in H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}$. Furthermore, $\forall \varepsilon>0, \forall g \in H^{1}\left(\mathbb{T}^{2}\right)$,

$$
\begin{align*}
I\left(A_{t}+\varepsilon g\right) & =\int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}}\left(A_{t}+\varepsilon g\right)-F_{t}\right|^{2} \varphi \\
& =\int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} A_{t}-F_{t}\right|^{2} \varphi-2 \varepsilon \int_{\mathbb{T}^{2}}\left(\nabla_{x_{1}^{2}} A_{t}-F_{t}\right) \cdot \nabla_{x_{1}^{2}} g \varphi+\varepsilon^{2} \int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} g\right|^{2} \varphi  \tag{3.7}\\
& =I\left(A_{t}\right)-2 \varepsilon \int_{\mathbb{T}^{2}}\left(\nabla_{x_{1}^{2}} A_{t}-F_{t}\right) \cdot \nabla_{x_{1}^{2}} g \varphi+\varepsilon^{2} \int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} g\right|^{2} \varphi
\end{align*}
$$

Since $A_{t}$ is the minimum of $I$, then $I\left(A_{t}+\varepsilon g\right) \geq I\left(A_{t}\right), \forall \varepsilon>0, \forall g \in H^{1}\left(\mathbb{T}^{2}\right)$. By considering the asymptotic regime $\varepsilon \rightarrow 0$ in the last equation, one thus obtains the equation (3.4):

$$
\int_{\mathbb{T}^{2}} \nabla_{x_{1}^{2}} A_{t} \cdot \nabla_{x_{1}^{2}} g \varphi=\int_{\mathbb{T}^{2}} F_{t} \cdot \nabla_{x_{1}^{2}} g \varphi, \forall g \in H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}
$$

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This is the weak formulation of the problem (3.5) in $H^{1}\left(\mathbb{T}^{2}\right) / \mathbb{R}$. Since $\varphi$ is a smooth positive function, then $\exists \delta>0$, s.t. $\varphi>\delta$. Furthermore, since $\operatorname{div}_{x_{1}^{2}}\left(F_{t} \varphi(t,).\right) \in L^{2}\left(\mathbb{T}^{2}\right)$, thus $\Delta_{x_{1}^{2}} A_{t} \in L^{2}\left(\mathbb{T}^{2}\right)$. Therefore, using standard elliptic regularity results, $A_{t} \in H^{2}\left(\mathbb{T}^{2}\right)$.

For any positive probability density function $\varphi$, the estimated vector field $\nabla_{x_{1}^{2}} A_{t}$ is the projection of $F_{t}$ onto a gradient. In the following, we will use the notation:

$$
\begin{equation*}
\mathcal{P}_{\varphi}\left(F_{t}\right)=\nabla_{x_{1}^{2}} A_{t} \tag{3.8}
\end{equation*}
$$

where the projection operator $\mathcal{P}_{\varphi}$ is a linear projection defined from $H_{\varphi}\left(\operatorname{div} ; \mathbb{T}^{2}\right)$ to $H^{1}\left(\mathbb{T}^{2}\right) \times H^{1}\left(\mathbb{T}^{2}\right)$. Notice in particular that $\mathcal{P}_{\varphi} \circ \mathcal{P}_{\varphi}=\mathcal{P}_{\varphi}$. Moreover, the projection operator $\mathcal{P}$ defined by (1.14) is $\mathcal{P}_{\varphi}$ for $\varphi=1$.

### 3.2. The projected ABF dynamics

We will study the longtime convergence of the following Projected ABF (PABF) dynamics:

$$
\left\{\begin{align*}
d X_{t} & =-\nabla\left(V-A_{t} \circ \xi\right)\left(X_{t}\right) d t+\sqrt{2 \beta^{-1}} d W_{t}  \tag{3.9}\\
\nabla_{x_{1}^{2}} A_{t} & =\mathcal{P}_{\psi^{\xi}}\left(F_{t}\right) \\
F_{t}^{i}\left(x_{1}, x_{2}\right) & =\mathbb{E}\left[\partial_{i} V\left(X_{t}\right) \mid \xi\left(X_{t}\right)=\left(x_{1}, x_{2}\right)\right], i=1,2
\end{align*}\right.
$$

where $\mathcal{P}_{\psi^{\xi}}$ is the linear projection defined by (3.8) and $W_{t}$ is a standard $n N$-dimensional Brownian motion. Thanks to the diffusion term $\sqrt{2 \beta^{-1}} d W_{t}, X_{t}$ admits a smooth density $\psi$ with respect to the the Lebesgue measure on $\mathbb{T}^{n}$ and $\psi^{\xi}$ then denotes the marginal distribution of $\psi$ along $\xi$ :

$$
\begin{equation*}
\psi^{\xi}\left(t, x_{1}, x_{2}\right)=\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \psi(t, x) d x_{3}^{n} \tag{3.10}
\end{equation*}
$$

The dynamics (3.9) is the PABF dynamics (1.16) with $\xi(x)=\left(x_{1}, x_{2}\right), W_{c}=0$ and the weighted Helmholtz projection $\mathcal{P}_{\psi^{\xi}}$. As already mentioned above, the weight $\psi^{\xi}$ is introduced to simplify the convergence proof (see Remark 4.1 in Section(4).

Remark 3.2. If the law of $X_{t}$ is $\psi(t, x) d x$ then the law of $\xi\left(X_{t}\right)$ is $\psi^{\xi}\left(t, x_{1}, x_{2}\right) d x_{1} d x_{2}$ and the conditional distribution of $X_{t}$ given $\xi\left(X_{t}\right)=\left(x_{1}, x_{2}\right)$ is (see (1.7) for a similar formula when $\psi=$ $\left.Z_{\Sigma_{\left(x_{1}, x_{2}\right)}}^{-1} \mathrm{e}^{-\beta V}\right):$

$$
\begin{equation*}
d \mu_{t, x_{1}, x_{2}}=\frac{\psi(t, x) d x_{3}^{n}}{\psi^{\xi}\left(t, x_{1}, x_{2}\right)} \tag{3.11}
\end{equation*}
$$

Indeed, for any smooth functions $f$ and $g$,

$$
\begin{aligned}
\mathbb{E}\left(f\left(\xi\left(X_{t}\right)\right) g\left(X_{t}\right)\right) & =\int_{\mathbb{T}^{n}} f(\xi(x)) g(x) \psi(t, x) d x \\
& =\int_{\mathbb{T}^{2}} \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} f \circ \xi g \psi d x_{3}^{n} d x_{1} d x_{2} \\
& =\int_{\mathbb{T}^{2}} f\left(x_{1}, x_{2}\right) \frac{\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} g \psi d x_{3}^{n}}{\psi^{\xi}\left(x_{1}, x_{2}\right)} \psi^{\xi}\left(x_{1}, x_{2}\right) d x_{1} d x_{2}
\end{aligned}
$$

$\diamond$
Let us now introduce the nonlinear partial differential equation (the so-called Fokker-Planck equation) which rules the evolution of the density $\psi(t, x)$ of $X_{t}$ solution of (3.9):

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$$
\left\{\begin{array}{l}
\partial_{t} \psi=\operatorname{div}\left(\left(\nabla V-\sum_{i=1}^{2} \partial_{i} A_{t} \circ \xi \nabla \xi_{i}\right) \psi+\beta^{-1} \nabla \psi\right), \quad \text { for }(t, x) \in\left[0, \infty\left[\times \mathbb{T}^{n}\right.\right.  \tag{3.12}\\
\forall t \geq 0, \operatorname{div}\left(\nabla A_{t} \psi^{\xi}(t, .)\right)=\operatorname{div}\left(F_{t} \psi^{\xi}(t, .)\right), \text { in } \mathbb{T}^{2} \text { with periodic boundary conditions, } \\
\forall t \geq 0, \forall\left(x_{1}, x_{2}\right) \in \mathbb{T}^{2}, F_{t}^{i}\left(x_{1}, x_{2}\right)=\frac{\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} V \psi d x_{3}^{n}}{\psi^{\xi}\left(x_{1}, x_{2}\right)}, i=1,2
\end{array}\right.
$$

The first equation of (3.12) rewrites:

$$
\begin{equation*}
\partial_{t} \psi=\operatorname{div}\left[\nabla V \psi+\beta^{-1} \nabla \psi\right]-\partial_{1}\left(\left(\partial_{1} A_{t}\right) \psi\right)-\partial_{2}\left(\left(\partial_{2} A_{t}\right) \psi\right) . \tag{3.13}
\end{equation*}
$$

Suppose that $\left(\psi, F_{t}, A_{t}\right)$ is a solution of (3.12) and let us introduce the expected long-time limits of $\psi, \psi^{\xi}$ (defined by (3.10)) and $\mu_{t, x_{1}, x_{2}}$ (defined by (3.11)) respectively:
(1) $\psi_{\infty}=\mathrm{e}^{-\beta(V-A \circ \xi)}$;
(2) $\psi_{\infty}^{\xi}=1$ (uniform law);
(3)

$$
\begin{equation*}
\mu_{\infty, x_{1}, x_{2}}=Z_{\Sigma_{\left(x_{1}, x_{2}\right)}}^{-1} \mathrm{e}^{-\beta V} d x_{3}^{n} . \tag{3.14}
\end{equation*}
$$

Notice that the probability measure $\psi_{\infty}^{\xi}\left(x_{1}, x_{2}\right) d x_{1} d x_{2}$ is the image of the probability measure $\psi_{\infty}(x) d x$ by $\xi$ and that $\mu_{\infty, x_{1}, x_{2}}=\mu_{\Sigma_{\left(x_{1}, x_{2}\right)}}$ defined in (1.7). Furthermore, we have that

$$
\int_{\mathbb{T}^{n}} \psi_{\infty}=1, \int_{\mathbb{T}^{2}} \psi_{\infty}^{\xi}=1 \text { and } \forall\left(x_{1}, x_{2}\right) \in \mathbb{T}^{2}, \quad \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} d \mu_{\infty, x_{1} x_{2}}=1
$$

Remark 3.3. In the case when $W_{c} \neq 0$, the first equation of the Fokker-Plank problem (3.12) becomes:

$$
\partial_{t} \psi=\operatorname{div}\left(\nabla\left(V-A_{t} \circ \xi-W_{c} \circ \xi\right) \psi+\beta^{-1} \nabla \psi\right) .
$$

The expected long-time limits of $\psi, \psi^{\xi}$ and $\mu_{t, x_{1}, x_{2}}$ are respectively:
(1) $\psi_{\infty}=Z_{W_{c}}^{-1} \mathrm{e}^{-\beta\left(V-A \circ \xi-W_{c} \circ \xi\right)}$;
(2) $\psi_{\infty}^{\xi}=Z_{W_{c}}^{-1} \mathrm{e}^{-\beta W_{c}}$;
(3) $\mu_{\infty, x_{1}, x_{2}}=Z_{\Sigma_{\left(x_{1}, x_{2}\right)}^{-1}} \mathrm{e}^{-\beta V} d x_{3}^{n}$,
where $Z_{W_{c}}=\int_{\mathbb{T}^{2}} \mathrm{e}^{-\beta W_{c}}$.

### 3.3. Precise statements of the longtime convergence results

In section 3.3.1, some well-known results on entropy techniques are presented. For a general introduction to logarithmic Sobolev inequalities, their properties and their relation to long-time behaviours of solutions to partial differential equations, we refer to $[1,2,13]$. Section 3.3 .2 presents the main theorem of convergence.

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### 3.3.1. Entropy and Fisher information

Define the relative entropy $H(. \mid$.$) as follows: for any probability measures \mu$ and $\nu$ such that $\mu$ is absolutely continuous with respect to $\nu$ (denoted $\mu \ll \nu$ ),

$$
H(\mu \mid \nu)=\int \ln \left(\frac{d \mu}{d \nu}\right) d \mu
$$

Abusing the notation, we will denote $H(\varphi \mid \psi)$ for $H(\varphi(x) d x \mid \psi(x) d x)$ in case of probability measures with densities. Let us recall the Csiszar-Kullback inequality (see [2]):

$$
\begin{equation*}
\|\mu-\nu\|_{T V} \leq \sqrt{2 H(\mu \mid \nu)} \tag{3.15}
\end{equation*}
$$

where $\|\mu-\nu\|_{T V}=\sup _{\|f\|_{L^{\infty} \leq 1}}\left\{\int f d(\mu-\nu)\right\}$ is the total variation norm of the signed measure $\mu-\nu$. When both $\mu$ and $\nu$ have densities with respect to the Lebesque measure, $\|\mu-\nu\|_{T V}$ is simply the $L^{1}$ norm of the difference between the two densities. The entropy $H(\mu \mid \nu)$ can be understood as a measure of how close $\mu$ and $\nu$ are.

Now, let us define the Fisher information of $\mu$ with respect to $\nu$ :

$$
\begin{equation*}
I(\mu \mid \nu)=\int\left|\nabla \ln \left(\frac{d \mu}{d \nu}\right)\right|^{2} d \mu \tag{3.16}
\end{equation*}
$$

The Wasserstein distance is another way to compare two probability measures $\mu$ and $\nu$ defined on a space $\Sigma$,

$$
\mathcal{W}(\mu, \nu)=\sqrt{\inf _{\pi \in \prod(\mu, \nu)} \int_{\Sigma \times \Sigma} d_{\Sigma}(x, y)^{2} d \pi(x, y)}
$$

where the geodesic distance $d_{\Sigma}$ on $\Sigma$ is defined as: $\forall x, y \in \Sigma$,

$$
d_{\Sigma}(x, y)=\inf \left\{\sqrt{\int_{0}^{1}|\dot{w}(t)|^{2} d t} \mid w \in C^{1}([0,1], \Sigma), w(0)=x, w(1)=y\right\}
$$

and $\Pi(\mu, \nu)$ denotes the set of coupling probability measures, namely probability measures on $\Sigma \times \Sigma$ such that their marginals are $\mu$ and $\nu$ :

$$
\forall \pi \in \prod(\mu, \nu), \int_{\Sigma \times \Sigma} \phi(x) d \pi(x, y)=\int_{\Sigma} \phi d \mu \text { and } \int_{\Sigma \times \Sigma} \psi(y) d \pi(x, y)=\int_{\Sigma} \psi d \nu
$$

Definition 3.4. We say that a probability measure $\nu$ satisfies a logarithmic Sobolev inequality with constant $\rho>0($ denoted $\operatorname{LSI}(\rho))$ if for all probability measure $\mu$ such that $\mu \ll \nu$,

$$
H(\mu \mid \nu) \leq \frac{1}{2 \rho} I(\mu \mid \nu)
$$

Definition 3.5. We say that a probability measure $\nu$ satisfies a Talagrand inequality with constant $\rho>0$ (denoted $\mathrm{T}(\rho))$ if for all probability measure $\mu$ such that $\mu \ll \nu$,

$$
\mathcal{W}(\mu, \nu) \leq \sqrt{\frac{2}{\rho} H(\mu \mid \nu)}
$$

Remark 3.6. We implicitly assume in the latter definition, that the probability measures have finite moments of order 2 . This is the case for the probability measures used in this paper.

The following lemma is proved in [11], Theorem 1:
Lemma 3.7. If $\nu$ satisfies $L S I(\rho)$, then $\nu$ satisfies $T(\rho)$.

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Recall that $X_{t}$ solution to (3.9) has a density $\psi(t,$.$) . In the following, we denote the Total Entropy$ by

$$
\begin{equation*}
E(t)=H\left(\psi(t, .) \mid \psi_{\infty}\right)=\int_{\mathbb{T}^{n}} \ln \left(\psi / \psi_{\infty}\right) \psi, \tag{3.17}
\end{equation*}
$$

the Macroscopic Entropy by

$$
\begin{equation*}
E_{M}(t)=H\left(\psi^{\xi}(t, .) \mid \psi_{\infty}^{\xi}\right)=\int_{\mathbb{T}^{2}} \ln \left(\psi^{\xi} / \psi_{\infty}^{\xi}\right) \psi^{\xi} \tag{3.18}
\end{equation*}
$$

and the Microscopic Entropy by

$$
\begin{equation*}
E_{m}(t)=\int_{\mathcal{M}} e_{m}\left(t, x_{1}, x_{2}\right) \psi^{\xi}\left(t, x_{1}, x_{2}\right) d x_{1} d x_{2} \tag{3.19}
\end{equation*}
$$

where $e_{m}\left(t, x_{1}, x_{2}\right)=H\left(\mu_{t, x_{1}, x_{2}} \mid \mu_{\infty, x_{1}, x_{2}}\right)$. The following result is straightforward to check:
Lemma 3.8. It holds, $\forall t \geq 0$,

$$
E(t)=E_{M}(t)+E_{m}(t)
$$

Note that the Fisher information of $\mu_{t, x_{1}, x_{2}}$ with respect to $\mu_{\infty, x_{1}, x_{2}}$ can be written as (see (3.16)):

$$
I\left(\mu_{t, x_{1}, x_{2}} \mid \mu_{\infty, x_{1}, x_{2}}\right)=\int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left|\nabla_{x_{3}^{n}} \ln \left(\psi(t, .) / \psi_{\infty}\right)\right|^{2} d \mu_{t, x_{1}, x_{2}}
$$

### 3.3.2. Convergence of the PABF dynamics (3.12)

The following proposition shows that the density function $\psi^{\xi}$ satisfies a simple diffusion equation.
Proposition 3.9. Suppose that $\left(\psi, F_{t}, A_{t}\right)$ is a smooth solution of (3.12). Then $\psi^{\xi}$ satisfies:

$$
\left\{\begin{align*}
\partial_{t} \psi^{\xi} & =\beta^{-1} \Delta_{x_{1}^{2}} \psi^{\xi}, \text { in }\left[0, \infty\left[\times \mathbb{T}^{2},\right.\right.  \tag{3.20}\\
\psi^{\xi}(0, .) & =\psi_{0}^{\xi}, \text { on } \mathbb{T}^{2}
\end{align*}\right.
$$

Remark 3.10. If $\psi_{0}^{\xi}=0$ at some points or is not smooth, then $F$ at time 0 may not be well defined or $I\left(\psi^{\xi}(0,.) / \psi_{\infty}^{\xi}\right)$ may be infinite. Since, by Proposition $3.9, \psi^{\xi}$ satisfies a simple diffusion equation these difficulties disappear as soon as $t>0$. Therefore, up to considering the problem for $t>t_{0}>0$, we can suppose that $\psi_{0}^{\xi}$ is a smooth positive function. We also have that for all $t>0, \psi^{\xi}(t,)>$.0 , $\int_{\mathbb{T}^{2}} \psi^{\xi}=1$ and $\psi^{\xi}(t,.) \in C^{\infty}\left(\mathbb{T}^{2}\right)$.
Remark 3.11. In the case where $W_{c} \neq 0$, the probability density function $\psi^{\xi}$ satisfies the modified diffusion equation:

$$
\partial_{t} \psi^{\xi}=\nabla_{x_{1}^{2}} \cdot\left(\beta^{-1} \nabla_{x_{1}^{2}} \psi^{\xi}+\psi^{\xi} \nabla_{x_{1}^{2}} W_{c}\right) .
$$

Here are two simple corollaries of Proposition 3.9.
Corollary 3.12. There exists $t_{0}>0$ and $I_{0}>0$ (depending on $\psi_{0}^{\xi}$ ), such that

$$
\forall t>t_{0}, \quad I\left(\psi^{\xi}(t, .) \mid \psi_{\infty}^{\xi}\right)<I_{0} \mathrm{e}^{-\beta^{-1} 8 \pi^{2} t}
$$

Corollary 3.13. The macroscopic entropy $E_{M}(t)$, defined by (3.18), converges exponentially fast to zero:

$$
\forall t>t_{0}, \quad E_{M}(t) \leq \frac{I_{0}}{8 \pi^{2}} \mathrm{e}^{-\beta^{-1} 8 \pi^{2} t}
$$

where $I_{0}$ is the constant introduced in Corollary 3.12.
The assumptions we need to prove the longtime convergence of the biasing force $\nabla A_{t}$ to the mean force $\nabla A$ are the following:

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[H1 ] $V \in C^{2}\left(\mathbb{T}^{n}\right)$ and satisfies:

$$
\exists \gamma>0, \forall 3 \leq j \leq n, \forall x \in \mathbb{T}^{n}, \quad \max \left(\left|\partial_{1} \partial_{j} V(x)\right|,\left|\partial_{2} \partial_{j} V(x)\right|\right) \leq \gamma
$$

[H2 ] $V$ is such that $\exists \rho>0, \forall\left(x_{1}, x_{2}\right) \in \mathbb{T}^{2}, \mu_{\infty, x_{1}, x_{2}}=\mu_{\Sigma\left(x_{1}, x_{2}\right)}$ defined by (3.14) satisfies LSI $(\rho)$.

The main theorem is:
Theorem 3.14. Let us assume [H1] and [H2]. The following properties then hold:
(1) The microscopic entropy $E_{m}$ converges exponentially fast to zero:

$$
\begin{equation*}
\exists C>0, \exists \lambda>0, \forall t \geq 0, \quad \sqrt{E_{m}(t)} \leq C \mathrm{e}^{-\lambda t} . \tag{3.21}
\end{equation*}
$$

Furthermore, if $\rho \neq 4 \pi^{2}$, then $\lambda=\beta^{-1} \min \left(\rho, 4 \pi^{2}\right)$ and
$C=\sqrt{E_{m}(0)}+\frac{\gamma}{\beta^{-1}\left|\rho-4 \pi^{2}\right|} \sqrt{\frac{I_{0}}{2 \rho}}$. If $\rho=4 \pi^{2}$, then for all $\lambda<\beta^{-1} \rho$, there exists a positive constant $C$ such that (3.21) is satisfied.
(2) $\sqrt{E(t)}$ and $\left\|\psi(t, .)-\psi_{\infty}\right\|_{L^{1}\left(\mathbb{T}^{n}\right)}$ both converge exponentially fast to zero with rate $\lambda$.
(3) The biasing force $\nabla_{x_{1}^{2}} A_{t}$ converges to the mean force $\nabla_{x_{1}^{2}} A$ in the following sense:

$$
\begin{equation*}
\forall t \geq 0, \int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} A_{t}-\nabla_{x_{1}^{2}} A\right|^{2} \psi^{\xi}\left(t, x_{1}, x_{2}\right) d x_{1} d x_{2} \leq \frac{8 \gamma^{2}}{\rho} E_{m}(t) \tag{3.22}
\end{equation*}
$$

The proofs of the results presented in this section are provided in Section 4.
Remark 3.15. We would like to emphasize that our arguments hold under the assumption of existence of regular solutions. In particular, we suppose that the density $\psi(t,$.$) is sufficiently regular so that the$ algebric manipulations in the proofs (see Section(4) are valid.
Remark 3.16. This remark is devoted to show how the rate of convergence of the original gradient dynamics (1.2) is improved thanks to PABF method. First of all, we mention a classical computation to get a rate of convergence for (1.2). Precisely, if one denotes $\varphi(t,$.$) the probability density function$ of $X_{t}$ satisfying (1.2), and $\varphi_{\infty}=Z_{\mu}^{-1} \mathrm{e}^{-\beta V}$ its longtime limit, then by standard computations (see for example [2]), one obtains:

$$
\frac{d}{d t} H\left(\varphi(t, .) \mid \varphi_{\infty}\right)=-\beta^{-1} I\left(\varphi(t, .) \mid \varphi_{\infty}\right) .
$$

Therefore, if $\varphi_{\infty}$ satisfies $\operatorname{LSI}(R)$, then one obtains the estimate

$$
\begin{equation*}
\exists R>0, \forall t>0, \quad H\left(\varphi(t, .) \mid \varphi_{\infty}\right) \leq H\left(\varphi_{0} \mid \varphi_{\infty}\right) \mathrm{e}^{-2 \beta^{-1} R t} . \tag{3.23}
\end{equation*}
$$

From (3.15), we obtain that $\left\|\varphi(t, .)-\varphi_{\infty}\right\|_{L^{1}\left(\mathbb{T}^{n}\right)}$ converges exponentially fast to zero with rate $\beta^{-1} R$. The constant $R$ is known to be small if the metastable states are separated by large energy barriers or if high probability regions for $\mu$ are separated by large regions with small probability (namely $\mu$ is a multimodal measure). Second, by Theorem 3.14, one can show that $\nabla A_{t}$ converges exponentially fast to $\nabla A$ in $L^{2}\left(\psi_{\infty}^{\xi}\left(x_{1}, x_{2}\right) d x_{1} d x_{2}\right)$-norm at rate $\lambda=\beta^{-1} \min \left(\rho, 4 \pi^{2}\right)$. Indeed, since $\psi_{\infty}^{\xi}=1$,

$$
\begin{aligned}
\int_{\mathbb{T}^{2}}\left|\nabla A_{t}-\nabla A\right|^{2} d x_{1} d x_{2} & =\int_{\mathbb{T}^{2}}\left|\nabla A_{t}-\nabla A\right|^{2} \frac{\psi^{\xi}\left(t, x_{1}, x_{2}\right)}{\psi^{\xi}\left(t, x_{1}, x_{2}\right)} d x_{1} d x_{2} \\
& \leq \frac{8 \gamma^{2}}{(1-\varepsilon) \rho} E_{m}(t), \\
& \leq \tilde{C} \mathrm{e}^{-2 \lambda t},
\end{aligned}
$$

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where $\varepsilon>0$ such that $\psi^{\xi}\left(t, x_{1}, x_{2}\right) \geq 1-\varepsilon$ (for more details refer to the proof of Corollary 3.12 in Section(4). This result can be compared with (3.23). Typically, for good choices of $\xi, \lambda \gg R$, the PABF dynamics converges to equilibrium much faster than the original dynamics (1.2). This is typically the case if the conditional measures $\mu_{\infty, x_{1}, x_{2}}$ are less multimodal than the original measure $\mu$. In our framework, we could state that a "good reaction coordinate" is such that $\rho$ is as large as possible.

Remark 3.17. (Extension to other geometric settings)
The results of Theorem 3.14 are easily generalized to the following setting:
If $\mathcal{D}=\mathbb{R}^{n}, \xi(x)=\left(x_{1}, x_{2}\right)$ and $\mathcal{M}$ is a compact subspace of $\mathbb{R}^{n}$, then choose a confining potential $W_{c}$ (defined in (1.11)) such that $Z_{W_{c}}=\int \mathrm{e}^{-\beta W_{c}}<+\infty, Z_{W_{c}}^{-1} \mathrm{e}^{-\beta W_{c}}$ satisfies $\operatorname{LSI}\left(r^{*}\right)$ (for some $r^{*}>0$ ) and $W_{c}$ is convex potential, then Corollary 3.12 is satisfied with rate $2 \beta^{-1}\left(r^{*}-\varepsilon\right)$, for any $\varepsilon \in\left(0, r^{*}\right)$ (refer to Corollary 1 in [9] for further details). In this case, Neumann boundary conditions are needed to solve the Poisson problem (3.5):

$$
\left\{\begin{align*}
\operatorname{div}\left(\nabla A_{t} \psi^{\xi}(t, .)\right) & =\operatorname{div}\left(F_{t} \psi^{\xi}(t, .)\right) & & \text { in } \mathcal{M}  \tag{3.24}\\
\frac{\partial A_{t}}{\partial n} & =F_{t} \cdot n & & \text { on } \partial \mathcal{M},
\end{align*}\right.
$$

where $n$ denotes the unit normal outward to $\mathcal{M}$. The convergence rate $\lambda$ of Theorem 3.14 becomes $\beta^{-1} \min \left(\rho, r^{*}-\varepsilon\right)$. Neumann boundary conditions come from the minimization problem (3.6) associated with the Euler-Lagrange equation. The numerical applications in Section 2 are performed in this setting.

Remark 3.18. (Extension to more general reaction coordinates)
In this section, we have chosen $\xi\left(x_{1}, \ldots, x_{n}\right)=\left(x_{1}, x_{2}\right)$. The results can be extended to the following settings:
(1) In dimension one, the Helmholtz projection has obviously no sense. However, if $\mathcal{D}=\mathbb{T}^{n}$ and $\xi(x)=x_{1}$, then $F_{t}$ converges to $A^{\prime}$, which is a derivative of a periodic function and thus $\int_{\mathbb{T}} A^{\prime}=0$. Since $\int_{\mathbb{T}} F_{t}$ is not necessary equal to zero, one can therefore take a new approximation $A_{t}^{\prime}=F_{t}-\int_{\mathbb{T}} F_{t}$, which approximates $A^{\prime}$ at any time $t$. The convergence results of this section can be extended to this setting, to show that $A_{t}^{\prime}$ converges exponentially fast to $A^{\prime}$.
(2) More generally, for a reaction coordinate with values in $\mathbb{T}^{m}$, the convergence results presented in this paper still hold under the following orthogonality condition:

$$
\begin{equation*}
\forall i \neq j, \nabla \xi_{i} \cdot \nabla \xi_{j}=0 \tag{3.25}
\end{equation*}
$$

The proof follows exactly the same lines. In the case when (3.25) does not hold, it is possible to resort to the following trick used for example in metadynamics (refer to $[3,8]$ ). The idea is to introduce an additional variable $z$ of dimension $m$, and an extended potential $V_{\xi}(x, z)=$ $V(x)+\frac{\kappa}{2}|z-\xi(x)|^{2}$, where $\kappa$ is a penalty constant. The reaction coordinate is then chosen as $\xi_{\text {meta }}(x, z)=z$, so that the associated free energy is:

$$
A_{\xi}(z)=-\beta^{-1} \ln \int_{\mathcal{D}} \mathrm{e}^{-\beta V_{\xi}(x, z)} d x
$$

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which converges to $A(z)$ (defined in (1.5)) when $\kappa$ goes to infinity. The extended PABF dynamics can be written as:

$$
\left\{\begin{aligned}
d X_{t} & =-\left(\nabla V\left(X_{t}\right)+\kappa \sum_{i=1}^{m}\left(\xi_{i}\left(X_{t}\right)-Z_{i, t}\right) \nabla \xi_{i}\left(X_{t}\right)\right) d t+\sqrt{2 \beta^{-1}} d W_{t}, \\
d Z_{t} & =\kappa\left(\xi\left(X_{t}\right)-\nabla E_{t}\left(Z_{t}\right)\right) d t+\sqrt{2 \beta^{-1}} d \bar{W}_{t}, \\
\nabla E_{t} & =\mathcal{P}_{\psi^{\xi_{m e t a}}}\left(G_{t}\right), \\
G_{t}(z) & =\mathbb{E}\left(\xi\left(X_{t}\right) \mid Z_{t}=z\right),
\end{aligned}\right.
$$

where $\bar{W}_{t}$ is a $m$-dimensional Brownian motion independent of $W_{t}$. The results of Theorem 3.14 apply to this extended PABF dynamics.

## 4. Proofs

The proofs are inspired from [9]. One may assume that $\beta=1$ up to the following change of variable: $\tilde{t}=\beta^{-1} t, \tilde{\psi}(\tilde{t}, x)=\psi(t, x), \tilde{V}(x)=\beta V(x)$. Recall, that we work in $\mathcal{D}=\mathbb{T}^{n}, \mathcal{M}=\mathbb{T}^{2}$ and $\forall x=$ $\left(x_{1}, \ldots, x_{n}\right) \in \mathbb{T}^{n}, \xi(x)=\left(x_{1}, x_{2}\right)$.

### 4.1. Proof of Proposition 3.9

Let $g: \mathbb{T}^{2} \rightarrow \mathbb{R}$ be a function in $H^{1}\left(\mathbb{T}^{2}\right)$.

$$
\begin{aligned}
\frac{d}{d t} \int_{\mathbb{T}^{2}} \psi^{\xi} g d x_{1} d x_{2}= & \frac{d}{d t} \int_{\mathbb{T}^{n}} \psi g \circ \xi d x_{1}^{n} \\
= & \int_{\mathbb{T}^{n}} \operatorname{div}\left[\left(\nabla V-\sum_{i=1}^{2} \partial_{i} A_{t} \circ \xi \nabla \xi_{i}\right) \psi+\nabla \psi\right] g \circ \xi d x_{1}^{n} \\
= & -\int_{\mathbb{T}^{n}} \sum_{j=1}^{2}\left[\left(\nabla V-\sum_{i=1}^{2} \partial_{i} A_{t} \circ \xi \nabla \xi_{i}\right) \psi+\nabla \psi\right] \cdot \nabla \xi_{j} \partial_{j} g \circ \xi d x_{1}^{n} \\
= & -\sum_{i=1}^{2} \int_{\mathbb{T}^{n}}\left[\left(\nabla V \cdot \nabla \xi_{i} \psi+\nabla \psi \cdot \nabla \xi_{i}\right] \partial_{i} g \circ \xi d x_{1}^{n}\right. \\
& +\sum_{i=1}^{2} \int_{\mathbb{T}^{n}} \partial_{i} A_{t} \circ \xi \psi \partial_{i} g \circ \xi d x_{1}^{n} .
\end{aligned}
$$

Applying Fubini's theorem, it holds:

$$
\begin{aligned}
\frac{d}{d t} \int_{\mathbb{T}^{2}} \psi^{\xi} g d x_{1} d x_{2}= & -\sum_{i=1}^{2} \int_{\mathbb{T}^{2}} \int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left[\partial_{i} V \psi+\partial_{i} \psi\right] d x_{3}^{n} \partial_{i} g\left(x_{1}, x_{2}\right) d x_{1} d x_{2} \\
& +\sum_{i=1}^{2} \int_{\mathbb{T}^{2}} \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} A_{t}\left(x_{1}, x_{2}\right) \psi d x_{3}^{n} \partial_{i} g\left(x_{1}, x_{2}\right) d x_{1} d x_{2} \\
= & -\sum_{i=1}^{2} \int_{\mathbb{T}^{2}} F_{t}^{i} \psi^{\xi} \partial_{i} g\left(x_{1}, x_{2}\right) d x_{1} d x_{2}-\sum_{i=1}^{2} \int_{\mathbb{T}^{2}} \partial_{i} \psi^{\xi} \partial_{i} g\left(x_{1}, x_{2}\right) d x_{1} d x_{2} \\
& +\sum_{i=1}^{2} \int_{\mathbb{T}^{2}} \partial_{i} A_{t}\left(x_{1}, x_{2}\right) \psi^{\xi} \partial_{i} g\left(x_{1}, x_{2}\right) d x_{1} d x_{2} \\
= & \int_{\mathbb{T}^{2}} \Delta \psi^{\xi} g\left(x_{1}, x_{2}\right) d x_{1} d x_{2}
\end{aligned}
$$

where we used (3.4) with $\varphi=\psi^{\xi}(t,$.$) . This is the weak formulation of:$

$$
\partial_{t} \psi^{\xi}=\Delta \psi^{\xi}, \text { on }\left[0, \infty\left[\times \mathbb{T}^{2} .\right.\right.
$$

Remark 4.1. The reason why we consider the weighted Helmholtz decomposition (3.3) with $\varphi=$ $\psi^{\xi}(t,$.$) in the PABF dynamics (3.9) instead of the standard one (1.15) is precisely to obtain this$ simple diffusion equation on the function $\psi^{\xi}$. This is will also be useful in the proof of Lemma 4.5 below.

### 4.2. Proof of Corollary $\mathbf{3 . 1 2}$

Let $\phi=\psi^{\xi}$ and $\phi_{\infty}=\psi_{\infty}^{\xi}=1$. It is known that $\forall t \geq 0$ and $\forall\left(x_{1}, x_{2}\right) \in \mathbb{T}^{2}, \phi$ satisfies:

$$
\begin{equation*}
\partial_{t} \phi=\Delta_{x_{1}^{2}} \phi \tag{4.1}
\end{equation*}
$$

Moreover (See Remark(3.10), it is assumed that and is such that

$$
\int_{\mathbb{T}^{2}} \phi(0, .)=1 \text { and } \phi(0, .) \geq 0 .
$$

Let us show that $\forall t \geq 0, \forall k>0,\|\phi(t, .)-1\|_{H^{k}\left(\mathbb{T}^{2}\right)} \leq\|\phi(0, .)-1\|_{H^{k}\left(\mathbb{T}^{2}\right)} \mathrm{e}^{-8 \pi^{2} t}$.
First, we prove that $\phi$ converges to 1 in $L^{2}\left(\mathbb{T}^{2}\right)$,

$$
\begin{aligned}
\frac{1}{2} \frac{d}{d t} \int_{\mathbb{T}^{2}}|\phi-1|^{2} & =\int_{\mathbb{T}^{2}} \partial_{t} \phi(\phi-1) \\
& =\int_{\mathbb{T}^{2}} \Delta \phi(\phi-1) \\
& =-\int_{\mathbb{T}^{2}} \nabla \phi \nabla(\phi-1) \\
& =-\int_{\mathbb{T}^{2}}|\nabla \phi|^{2} \\
& \leq-4 \pi^{2} \int_{\mathbb{T}^{2}}|\phi-1|^{2}
\end{aligned}
$$

where we have used the Poincaré-Wirtinger inequality on the torus $\mathbb{T}^{2}$, applied to $\phi$ : for any function $f \in H^{1}\left(\mathbb{T}^{2}\right)$,

$$
\int_{\mathbb{T}^{2}}\left(f-\int_{\mathbb{T}^{2}} f\right)^{2} \leq \frac{1}{4 \pi^{2}} \int_{\mathbb{T}^{2}}|\nabla f|^{2}
$$

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We therefore obtain, $\|\phi(t, .)-1\|_{L^{2}\left(\mathbb{T}^{2}\right)}^{2} \leq\|\phi(0, .)-1\|_{L^{2}\left(\mathbb{T}^{2}\right)}^{2} \mathrm{e}^{-8 \pi^{2} t}$.
Second, we prove that $\partial_{i} \phi$ converges to 0 in $L^{2}\left(\mathbb{T}^{2}\right)$. For $i=1,2, \partial_{i} \phi$ satisfies (4.1): $\partial_{t}\left(\partial_{i} \phi\right)=$ $\Delta_{x_{1}^{2}}\left(\partial_{i} \phi\right)$, with periodic boundary conditions. As above,

$$
\begin{aligned}
\frac{1}{2} \frac{d}{d t} \int_{\mathbb{T}^{2}}\left|\partial_{i} \phi\right|^{2} & =\int_{\mathbb{T}^{2}} \partial_{t}\left(\partial_{i} \phi\right) \partial_{i} \phi \\
& =\int_{\mathbb{T}^{2}} \Delta\left(\partial_{i} \phi\right) \partial_{i} \phi \\
& =-\int_{\mathbb{T}^{2}}\left|\nabla\left(\partial_{i} \phi\right)\right|^{2} .
\end{aligned}
$$

Using again the Poincaré-Wirtinger inequality on $\partial_{i} \phi$,

$$
\begin{aligned}
\frac{1}{2} \frac{d}{d t} \int_{\mathbb{T}^{2}}\left|\partial_{i} \phi\right|^{2} & \leq-4 \pi^{2} \int_{\mathbb{T}^{2}}\left(\partial_{i} \phi-\int_{\mathbb{T}^{2}} \partial_{i} \phi\right)^{2} \\
& =-4 \pi^{2} \int_{\mathbb{T}^{2}}\left|\partial_{i} \phi\right|^{2} .
\end{aligned}
$$

Where we used $\int_{\mathbb{T}^{2}} \partial_{i} \phi=0$, since $\phi$ is periodic on $\mathbb{T}^{2}$. Therefore, it holds

$$
\left\|\partial_{i} \phi(t, .)\right\|_{L^{2}\left(\mathbb{T}^{2}\right)}^{2} \leq\left\|\partial_{i} \phi(0, .)\right\|_{L^{2}\left(\mathbb{T}^{2}\right)}^{2} \mathrm{e}^{-8 \pi^{2} t} .
$$

Third, one can prove by induction that all higher derivatives of $\phi$ converge exponentially fast to 0 , with rate $8 \pi^{2}$ and the following estimation is then proven:

$$
\forall t \geq 0, \forall k>0, \quad\|\phi(t, .)-1\|_{H^{k}\left(\mathbb{T}^{2}\right)}^{2} \leq\|\phi(0, .)-1\|_{H^{k}\left(\mathbb{T}^{2}\right)}^{2} \mathrm{e}^{-8 \pi^{2} t}
$$

As $H^{k}\left(\mathbb{T}^{2}\right) \hookrightarrow L^{\infty}\left(\mathbb{T}^{2}\right), \forall k>1$, then $\exists c>0$,

$$
\|\phi-1\|_{L^{\infty}}^{2} \leq c\|\phi-1\|_{H^{k}}^{2} \leq c \mathrm{e}^{-8 \pi^{2} t}
$$

Therefore, $\forall \varepsilon>0, \exists t_{0}>0, \forall x \in \mathbb{T}^{2}, \forall t>t_{0}, \phi(t, x) \geq 1-\varepsilon$. Finally, $\forall t>t_{0}$
$I\left(\psi^{\xi} \mid \psi_{\infty}^{\xi}\right)=\int_{\mathbb{T}^{2}} \frac{\left|\nabla_{x_{1}^{2}} \phi\right|^{2}}{\phi} \leq \frac{1}{1-\varepsilon} \int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} \phi\right|^{2} \leq \frac{\left\|\nabla_{x_{1}^{2}} \phi(0, .)\right\|_{L^{2}\left(\mathbb{T}^{2}\right)}^{2}}{1-\varepsilon} \mathrm{e}^{-8 \pi^{2} t}$.

### 4.3. Proof of Corollary $\mathbf{3 . 1 3}$

We have that $\psi_{\infty}^{\xi}=1$ satisfies $\operatorname{LSI}(r)$, for some $r>0$ (see Chapter 3, Section 3 in [1]). Referring to Proposition 3.9 and since $\psi^{\xi}$ is a probability density function, one gets:

$$
\begin{aligned}
\frac{d}{d t} E_{M} & =\int_{\mathbb{T}^{2}} \partial_{t}\left(\psi^{\xi} \ln \left(\psi^{\xi}\right)\right) \\
& =\int_{\mathbb{T}^{2}} \partial_{t} \psi^{\xi} \ln \left(\psi^{\xi}\right)+\int_{\mathbb{T}^{2}} \partial_{t} \psi^{\xi} \\
& =\int_{\mathbb{T}^{2}} \Delta \psi^{\xi} \ln \left(\psi^{\xi}\right) \\
& =-\int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} \ln \left(\psi^{\xi}\right)\right|^{2} \psi^{\xi} \\
& =-I\left(\psi^{\xi} \mid \psi_{\infty}^{\xi}\right) \\
& \leq-2 r H\left(\psi^{\xi} \mid \psi_{\infty}^{\xi}\right) \\
& =-2 r E_{M} .
\end{aligned}
$$

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Therefore, $E_{M}$ converges exponentially fast to zero. Referring to Corollary 3.12 and since $E_{M}$ converges to zero, we have that for any $t>t_{0}$,

$$
\begin{aligned}
-E_{M}(t)=\int_{t}^{\infty} \frac{d}{d s} E_{M}(s) d s & =-\int_{t}^{\infty} I\left(\psi^{\xi} \mid \psi_{\infty}^{\xi}\right) d s \\
& \geq-I_{0} \int_{t}^{\infty} \mathrm{e}^{-8 \pi^{2} s} d s \\
& =-\frac{I_{0}}{8 \pi^{2}} \mathrm{e}^{-8 \pi^{2} t}
\end{aligned}
$$

which yields the desired estimate.

### 4.4. Proof of Theorem $\mathbf{3 . 1 4}$

To prove our main result, several intermediate lemmas are needed.
Lemma 4.2. $\forall t \geq 0, \forall\left(x_{1}, x_{2}\right) \in \mathbb{T}^{2}$ and for $i=1,2$, we have:

$$
\left(F_{t}^{i}-\partial_{i} A\right)\left(x_{1}, x_{2}\right)=\left(\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} \ln \left(\psi / \psi_{\infty}\right) \frac{\psi}{\psi^{\xi}} d x_{3}^{n}\right)\left(x_{1}, x_{2}\right)-\left(\partial_{i} \ln \left(\psi^{\xi} / \psi_{\infty}^{\xi}\right)\right)\left(x_{1}, x_{2}\right)
$$

## Proof.

$$
\begin{aligned}
& \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} \ln \left(\psi / \psi_{\infty}\right) \frac{\psi}{\psi^{\xi}} d x_{3}^{n}-\partial_{i} \ln \left(\psi^{\xi} / \psi_{\infty}^{\xi}\right) \\
& =\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} \ln (\psi) \frac{\psi}{\psi^{\xi}} d x_{3}^{n}-\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} \ln \left(\psi_{\infty}\right) \frac{\psi}{\psi^{\xi}} d x_{3}^{n}-\partial_{i} \ln \left(\psi^{\xi}\right)+\partial_{i} \ln \left(\psi_{\infty}^{\xi}\right) \\
& =\frac{1}{\psi^{\xi}} \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{i} \psi d x_{3}^{n}+\int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left(\partial_{i} V-\nabla \xi_{i} \partial_{i} A \circ \xi\right) \frac{\psi}{\psi^{\xi}} d x_{3}^{n}-\partial_{i} \ln \left(\psi^{\xi}\right) \\
& =\frac{\partial_{i} \psi^{\xi}}{\psi^{\xi}}+F_{t}^{i}-\partial_{i} A-\frac{\partial_{i} \psi^{\xi}}{\psi^{\xi}} \\
& =F_{t}^{i}-\partial_{i} A
\end{aligned}
$$

Lemma 4.3. Suppose that [H1] and [H2] hold, then for all $t \geq 0$, for all $\left(x_{1}, x_{2}\right) \in \mathbb{T}^{2}$ and for $i=1,2$, we have:

$$
\left|F_{t}^{i}\left(x_{1}, x_{2}\right)-\partial_{i} A\left(x_{1}, x_{2}\right)\right| \leq \gamma \sqrt{\frac{2}{\rho} e_{m}\left(t, x_{1}, x_{2}\right)}
$$

Proof. For any coupling measure $\pi \in \prod\left(\mu_{t, x_{1}, x_{2}}, \mu_{\infty, x_{1}, x_{2}}\right)$ defined on $\Sigma_{\left(x_{1}, x_{2}\right)} \times \Sigma_{\left(x_{1}, x_{2}\right)}$, it holds:

$$
\begin{aligned}
\left|F_{t}^{i}-\partial_{i} A\right| & =\left|\int_{\Sigma_{\left(x_{1}, x_{2}\right)} \times \Sigma_{\left(x_{1}, x_{2}\right)}}\left(\partial_{i} V(x)-\partial_{i} V\left(x^{\prime}\right)\right) \pi\left(d x, d x^{\prime}\right)\right| \\
& =\left\|\nabla_{x_{3}^{n}} \partial_{i} V\right\|_{L^{\infty}} \sqrt{\int_{\Sigma_{\left(x_{1}, x_{2}\right)} \times \Sigma_{\left(x_{1}, x_{2}\right)}} d_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left(x, x^{\prime}\right)^{2} \pi\left(d x, d x^{\prime}\right)}
\end{aligned}
$$

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Taking now the infimum over all $\pi \in \prod\left(\mu\left(t, . \mid\left(x_{1}, x_{2}\right)\right), \mu\left(\infty, . \mid\left(x_{1}, x_{2}\right)\right)\right)$ and using Lemma 3.7, we obtain

$$
\begin{aligned}
\left|F_{t}^{i}-\partial_{i} A\right| & \leq \gamma \mathcal{W}\left(\mu\left(t, \mid\left(x_{1}, x_{2}\right)\right), \mu^{\xi}\left(\infty, \mid\left(x_{1}, x_{2}\right)\right)\right) \\
& \leq \gamma \sqrt{\frac{2}{\rho} H\left(\mu^{\xi}\left(t, \mid\left(x_{1}, x_{2}\right)\right), \mu^{\xi}\left(\infty, . \mid\left(x_{1}, x_{2}\right)\right)\right)} \\
& =\gamma \sqrt{\frac{2}{\rho} e_{m}\left(t,\left(x_{1}, x_{2}\right)\right)}
\end{aligned}
$$

Lemma 4.4. Suppose that [H2] holds, then for all $t \geq 0$,

$$
E_{m}(t) \leq \frac{1}{2 \rho} \int_{\mathbb{T}^{n}}\left|\nabla_{x_{3}^{n}} \ln \left(\psi(t, .) / \psi_{\infty}\right)\right|^{2} \psi .
$$

Proof. Using [H2],

$$
\begin{aligned}
E_{m} & =\int_{\mathbb{T}^{2}} e_{m} \psi^{\xi} d x_{1} d x_{2} \\
& =\int_{\mathbb{T}^{2}} H\left(\mu\left(t, . \mid\left(x_{1}, x_{2}\right)\right) \mid \mu\left(\infty, \mid\left(x_{1}, x_{2}\right)\right)\right) \psi^{\xi} d x_{1} d x_{2} \\
& \leq \int_{\mathbb{T}^{2}} \frac{1}{2 \rho} \int_{\Sigma_{x_{1}^{2}}}\left|\nabla_{x_{3}^{n}} \ln \left(\psi(t, .) / \psi_{\infty}\right)\right|^{2} d x_{3}^{n} \frac{\psi(t, .)}{\psi^{\xi}\left(t, x_{1}, x_{2}\right)} d x_{1} d x_{2} \\
& =\frac{1}{2 \rho} \int_{\mathbb{T}^{n}}\left|\nabla_{x_{3}^{n}} \ln \left(\psi(t, .) / \psi_{\infty}\right)\right|^{2} \psi d x_{1}^{n} .
\end{aligned}
$$

Lemma 4.5. It holds for all $t \geq 0$,

$$
\int_{\mathbb{T}^{n}}\left(\partial_{1} A_{t}-F_{t}^{1}\right)\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi+\int_{\mathbb{T}^{n}}\left(\partial_{2} A_{t}-F_{t}^{2}\right)\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi \leq 0 .
$$

Proof. Using Fubini's theorem,

$$
\begin{aligned}
& \int_{\mathbb{T}^{n}}\left(\partial_{1} A_{t}-F_{t}^{1}\right)\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi+\int_{\mathbb{T}^{n}}\left(\partial_{2} A_{t}-F_{t}^{2}\right)\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi \\
& =\int_{\mathbb{T}^{2}}\left(\partial_{1} A_{t}-F_{t}^{1}\right) \int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi+\int_{\mathbb{T}^{2}}\left(\partial_{2} A_{t}-F_{t}^{2}\right) \int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi
\end{aligned}
$$

For the first term, we have

$$
\begin{aligned}
\int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi & =\int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left(\partial_{1} \ln \psi\right) \psi-\int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left(\partial_{1} \ln \psi_{\infty}\right) \psi \\
& =\partial_{1} \psi^{\xi}+\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{1}(V-A) \psi \\
& =\left(\partial_{1} \ln \psi^{\xi}\right) \psi^{\xi}+F_{t}^{1} \psi^{\xi}-\partial_{1} A \psi^{\xi}
\end{aligned}
$$

Similarly, we have

$$
\int_{\Sigma_{\left(x_{1}, x_{2}\right)}}\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi=\left(\partial_{2} \ln \psi^{\xi}\right) \psi^{\xi}+F_{t}^{2} \psi^{\xi}-\partial_{2} A \psi^{\xi}
$$

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Therefore, one gets

$$
\begin{aligned}
& \int_{\mathbb{T}^{n}}\left(\partial_{1} A_{t}-F_{t}^{1}\right)\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi+\int_{\mathbb{T}^{n}}\left(\partial_{2} A_{t}-F_{t}^{2}\right)\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi \\
& =\int_{\mathbb{T}^{2}}\left(\partial_{1} A_{t}-F_{t}^{1}\right)\left(\partial_{1} \ln \psi^{\xi}\right) \psi^{\xi}+\int_{\mathbb{T}^{2}}\left(\partial_{2} A_{t}-F_{t}^{2}\right)\left(\partial_{2} \ln \psi^{\xi}\right) \psi^{\xi} \\
& \quad-\int_{\mathbb{T}^{2}}\left(\partial_{1} A_{t}-F_{t}^{1}\right)^{2} \psi^{\xi}-\int_{\mathbb{T}^{2}}\left(\partial_{2} A_{t}-F_{t}^{2}\right)^{2} \psi^{\xi}
\end{aligned}
$$

which concludes the assertion since the first line is equal to zero (by (3.4) with $\varphi=\psi^{\xi}(t,$.$) ) and the$ second line is non positive. Again the weighted Helmholtz decomposition helps in simplifying terms.

Proof of Theorem 3.14. We will now prove the exponentially convergence of $E_{m}(t)$ to zero. Recall (3.13):

$$
\partial_{t} \psi=\operatorname{div}(\nabla V \psi+\nabla \psi)-\partial_{1}\left(\left(\partial_{1} A_{t}\right) \psi\right)-\partial_{2}\left(\left(\partial_{2} A_{t}\right) \psi\right),
$$

which is equivalent to

$$
\partial_{t} \psi=\operatorname{div}\left(\psi_{\infty} \nabla\left(\frac{\psi}{\psi_{\infty}}\right)\right)+\partial_{1}\left[\left(\partial_{1} A-\partial_{1} A_{t}\right) \psi\right]+\partial_{2}\left[\left(\partial_{2} A-\partial_{2} A_{t}\right) \psi\right] .
$$

Using (3.17), (3.18) and (4.1), one obtains

$$
\begin{aligned}
\frac{d E}{d t}= & -\int_{\mathbb{T}^{n}}\left|\nabla \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi+\int_{\mathbb{T}^{n}}\left(\partial_{1} A_{t}-\partial_{1} A\right)\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi \\
& +\int_{\mathbb{T}^{n}}\left(\partial_{2} A_{t}-\partial_{2} A\right)\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi, \\
\frac{d E_{M}}{d t}= & -\int_{\mathbb{T}^{2}}\left|\nabla \ln \left(\psi^{\xi}\right)\right|^{2} \psi^{\xi} .
\end{aligned}
$$

Using then Lemma 3.8 and Lemma 4.2, one gets

$$
\begin{aligned}
\frac{d E_{m}}{d t}= & \frac{d E}{d t}-\frac{d E_{M}}{d t} \\
= & -\int_{\mathbb{T}^{n}}\left|\nabla \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi+\int_{\mathbb{T}^{n}}\left(\partial_{1} A_{t}-\partial_{1} A\right) \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{n}}\left(\partial_{2} A_{t}-\partial_{2} A\right) \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi+\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \psi^{\xi}\right|^{2} \psi^{\xi}+\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \psi^{\xi}\right|^{2} \psi^{\xi} \\
= & -\int_{\mathbb{T}^{n}}\left|\nabla \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi \\
& +\int_{\mathbb{T}^{n}}\left(\partial_{1} A_{t}-F_{t}^{1}\right)\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi+\int_{\mathbb{T}^{n}}\left(F_{t}^{1}-\partial_{1} A\right)\left[\partial_{1} \ln \left(\psi / \psi_{\infty}\right)\right] \psi \\
& +\int_{\mathbb{T}^{n}}\left(\partial_{2} A_{t}-F_{t}^{2}\right)\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi+\int_{\mathbb{T}^{n}}\left(F_{t}^{2}-\partial_{2} A\right)\left[\partial_{2} \ln \left(\psi / \psi_{\infty}\right)\right] \psi \\
& +\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \psi^{\xi}\right|^{2} \psi^{\xi}+\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \psi^{\xi}\right|^{2} \psi^{\xi} .
\end{aligned}
$$

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Lemma 4.5 then yields

$$
\begin{aligned}
\frac{d E_{m}}{d t} \leq & -\int_{\mathbb{T}^{n}}\left|\nabla \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi \\
& +\int_{\mathbb{T}^{n}}\left(F_{t}^{1}-\partial_{1} A\right) \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi+\int_{\mathbb{T}^{n}}\left(F_{t}^{2}-\partial_{2} A\right) \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \psi^{\xi}\right|^{2} \psi^{\xi}+\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \psi^{\xi}\right|^{2} \psi^{\xi}
\end{aligned}
$$

Using lemma 4.2 and Fubini's theorem, one then obtains

$$
\begin{aligned}
\frac{d E_{m}}{d t} \leq & -\int_{\mathbb{T}^{n}}\left|\nabla \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi \\
& +\int_{\mathbb{T}^{2}}\left[\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \frac{\psi}{\psi^{\xi}}\right] \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi-\int_{\mathbb{T}^{n}} \partial_{1} \ln \left(\psi^{\xi}\right) \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{2}}\left[\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \frac{\psi}{\psi^{\xi}}\right] \int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi-\int_{\mathbb{T}^{n}} \partial_{2} \ln \left(\psi^{\xi}\right) \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \psi^{\xi}\right|^{2} \psi^{\xi}+\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \psi^{\xi}\right|^{2} \psi^{\xi} \\
\leq & -\int_{\mathbb{T}^{n}}\left|\nabla \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi \\
& +\int_{\mathbb{T}^{2}}\left[\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi\right]^{2} \frac{1}{\psi^{\xi}}-\int_{\mathbb{T}^{n}} \partial_{1} \ln \left(\psi^{\xi}\right) \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{2}}\left[\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi\right]^{2} \frac{1}{\psi^{\xi}}-\int_{\mathbb{T}^{n}} \partial_{2} \ln \left(\psi^{\xi}\right) \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \psi^{\xi}\right|^{2} \psi^{\xi}+\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \psi^{\xi}\right|^{2} \psi^{\xi} .
\end{aligned}
$$

Applying the Cauchy-Schwarz inequality on the first terms of the second and third lines, we obtain

$$
\begin{aligned}
\frac{d E_{m}}{d t} \leq & -\int_{\mathbb{T}^{n}}\left|\nabla_{x_{3}^{n}} \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi \\
& -\int_{\mathbb{T}^{n}} \partial_{1} \ln \left(\psi^{\xi}\right) \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \psi-\int_{\mathbb{T}^{n}} \partial_{2} \ln \left(\psi^{\xi}\right) \partial_{2} \ln \left(\psi / \psi_{\infty}\right) \psi \\
& +\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \psi^{\xi}\right|^{2} \psi^{\xi}+\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \psi^{\xi}\right|^{2} \psi^{\xi} \\
\leq & -\int_{\mathbb{T}^{n}}\left|\nabla_{x_{3}^{n}} \ln \left(\psi / \psi_{\infty}\right)\right|^{2} \psi-\int_{\mathbb{T}^{2}} \partial_{1} \ln \left(\psi^{\xi}\right)\left[\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{1} \ln \left(\psi / \psi_{\infty}\right) \frac{\psi}{\psi^{\xi}}-\partial_{1} \ln \psi^{\xi}\right] \psi^{\xi} \\
& -\int_{\mathbb{T}^{2}} \partial_{2} \ln \left(\psi^{\xi}\right)\left[\int_{\Sigma_{\left(x_{1}, x_{2}\right)}} \partial_{2} \ln \left(\psi / \psi_{\infty} \frac{\psi}{\psi^{\xi}}-\partial_{2} \ln \psi^{\xi}\right] \psi^{\xi} .\right.
\end{aligned}
$$

Applying Lemma 4.4, Lemma 4.2, the Cauchy-Schwarz inequality, Lemma 4.3 and Corollary 3.12,

$$
\begin{aligned}
\frac{d E_{m}}{d t} \leq & -2 \rho E_{m}+\sqrt{\int_{\mathbb{T}^{2}}\left|\partial_{1} \ln \left(\psi^{\xi}\right)\right|^{2} \psi^{\xi}} \sqrt{\int_{\mathbb{T}^{2}} \frac{2}{\rho} e_{m}\left(t,\left(x_{1}, x_{2}\right)\right) \psi^{\xi}} \\
& +\sqrt{\int_{\mathbb{T}^{2}}\left|\partial_{2} \ln \left(\psi^{\xi}\right)\right|^{2} \psi^{\xi}} \sqrt{\int_{\mathbb{T}^{2}} \frac{2}{\rho} e_{m}\left(t,\left(x_{1}, x_{2}\right)\right) \psi^{\xi}} \\
\leq & -2 \rho E_{m}+2 \gamma \sqrt{\frac{2}{\rho} E_{m}} \sqrt{\int_{\mathbb{T}^{2}}\left|\nabla_{x_{1}^{2}} \ln \left(\psi^{\xi}\right)\right|^{2} \psi^{\xi}} \\
\leq & -2 \rho E_{m}+2 \gamma \sqrt{\frac{2}{\rho} E_{m}} \sqrt{I\left(\psi^{\xi} / \psi_{\infty}^{\xi}\right)} \\
\leq & -2 \rho E_{m}+2 \gamma \sqrt{\frac{2}{\rho} E_{m}} \sqrt{I_{0}} \mathrm{e}^{-4 \pi^{2} t} .
\end{aligned}
$$

Finally we obtain

$$
\frac{d}{d t} \sqrt{E_{m}(t)} \leq-\rho \sqrt{E_{m}(t)}+\gamma \sqrt{\frac{I_{0}}{2 \rho}} \mathrm{e}^{-4 \pi^{2} t}
$$

First, if $\rho \neq 4 \pi^{2}$, using Grönwall's inequality, one obtains

$$
\begin{aligned}
\sqrt{E_{m}(t)} & \leq \sqrt{E_{m}(0)} \mathrm{e}^{-\rho t}+\gamma \sqrt{\frac{I_{0}}{2 \rho}} \int_{0}^{t} \mathrm{e}^{\rho(-t+s)} \mathrm{e}^{-4 \pi^{2} s} d s \\
& \leq \sqrt{E_{m}(0)} \mathrm{e}^{-\rho t}+\gamma \sqrt{\frac{I_{0}}{2 \rho}} \frac{\mathrm{e}^{-\rho t}}{\rho-4 \pi^{2}}\left(\mathrm{e}^{\left(\rho-4 \pi^{2}\right) t}-1\right) \\
& \leq \sqrt{E_{m}(0)} \mathrm{e}^{-\rho t}+\gamma \sqrt{\frac{I_{0}}{2 \rho}} \frac{\mathrm{e}^{-\rho t}}{\left|\rho-4 \pi^{2}\right|} \mathrm{e}^{\left(\rho-4 \pi^{2}\right) t}
\end{aligned}
$$

Second, if $\rho=4 \pi^{2}$, one has

$$
\begin{aligned}
\sqrt{E_{m}(t)} & \leq \sqrt{E_{m}(0)} \mathrm{e}^{-\rho t}+\gamma \sqrt{\frac{I_{0}}{2 \rho}} \int_{0}^{t} \mathrm{e}^{-\rho t} d s \\
& \leq\left(\sqrt{E_{m}(0)}+\gamma \sqrt{\frac{I_{0}}{2 \rho}} t\right) \mathrm{e}^{-\rho t},
\end{aligned}
$$

which leads the desired estimation (3.21).
Using this convergence result, Corollary 3.13 and Lemma 3.8, it is then easy to see that $E$ converges exponentially fast to zero. Using (3.15), one obtains the convergence of $\psi$ to $\psi_{\infty}$ since:

$$
\left\|\psi-\psi_{\infty}\right\|_{L^{1}\left(\mathbb{T}^{n}\right)} \leq \sqrt{2 H\left(\psi \mid \psi_{\infty}\right)}=\sqrt{2 E}
$$

The second point of the theorem is checked. Finally, we are now in position to prove the last point of Theorem 3.14. Using (3.6) and Lemma 4.3,

$$
\begin{aligned}
\left\|\nabla A_{t}-\nabla A\right\|_{L_{\psi \xi}^{2}\left(\mathbb{T}^{2}\right)}^{2} & \leq 2\left\|\nabla A_{t}-F_{t}\right\|_{L_{\psi \xi}^{2}\left(\mathbb{T}^{2}\right)}^{2}+2\left\|F_{t}-\nabla A\right\|_{L_{\psi}{ }^{\xi}\left(\mathbb{T}^{2}\right)}^{2} \\
& \leq 4\left\|F_{t}-\nabla A\right\|_{L_{\psi}^{2}}^{2}\left(\mathbb{T}^{2}\right) \\
& \leq 8 \frac{\gamma^{2}}{\rho} E_{m} .
\end{aligned}
$$

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